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Iowa State University, Ph.D., 1972
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Determination of the Fermi surface of molybdenum
using the de Haas-van Alphen effect

by

James Alvin Hoekstra

A Dissertation Submitted to the
Graduate Faculty in Partial Fulfillment of
The Requirements for the Degree of
DOCTOR OF PHILOSOPHY

Department: Physics
Major: Solid State Physics

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1. INTRODUCTION

The Fermi surface (FS) and the electronic structure of molybdenum have been the subjects of several extensive investigations over the past few years. Most of the results have been interpreted in terms of revisions of a model for the chromium group metals (Cr, Mo, W) first proposed by Lomer (1), who used the energy bands for Fe as calculated by Wood (2). The initial model was later corrected for the particular case of molybdenum (3). Theoretical augmented-plane-wave (APW) calculations for Mo done by Loucks (4) exhibit qualitative agreement with the Lomer model, but the quantitative results are subject to numerical errors due to an incorrect size for the unit cell used in determining the potential applied in the calculation (5).

The rhombic dodecahedron which defines the Brillouin zone, shown in Figure 1, is labelled in accordance with the standard notation (6). As can be seen from the Lomer model in Figures 2 and 3, the FS consists of several distinct pieces. The largest two are the hole "octahedron" situated at the symmetry point H and the electron "jack" located at Γ . The "jack" consists of an octahedrally-shaped "body" with a "ball" protruding from each of the six corners of the "body." The narrow region connecting the "body" with one of the "balls" is referred to as a "neck." A hole "ellipsoid" is located at each of the symmetry points N. Six small electron "lenses" are situated along the lines Γ H in the "neck" regions of the "jack." In the remainder of this paper the terms used to describe these FS pieces will no longer be enclosed in quotation marks.

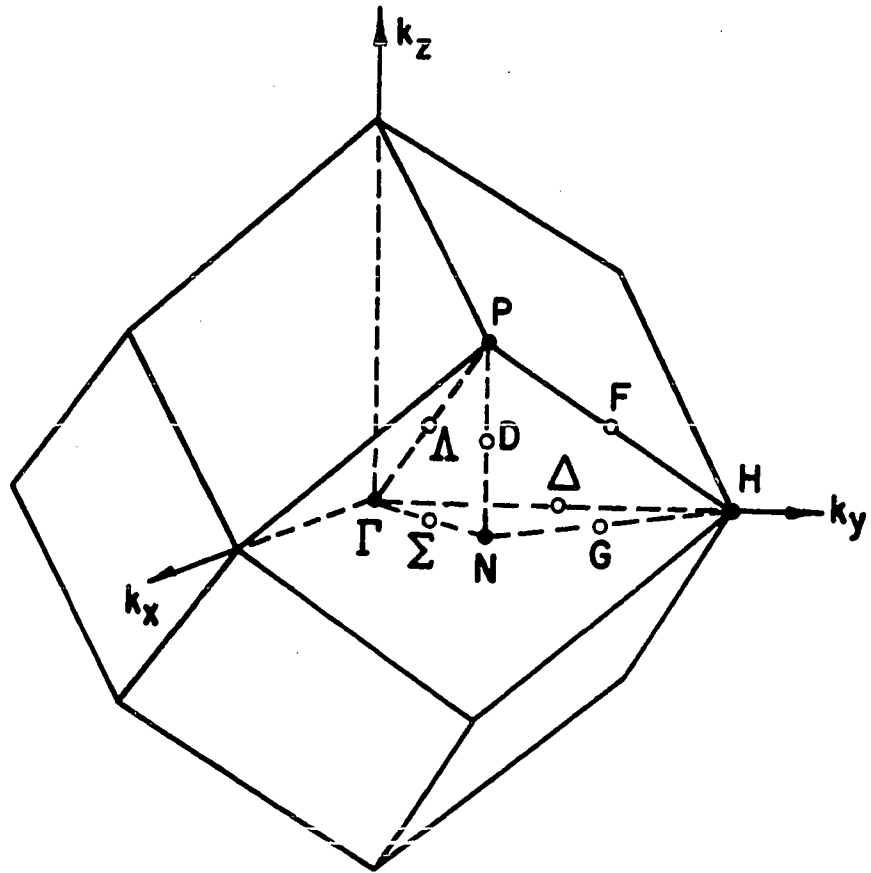


Figure 1. Brillouin zone for body-centered-cubic crystal lattice

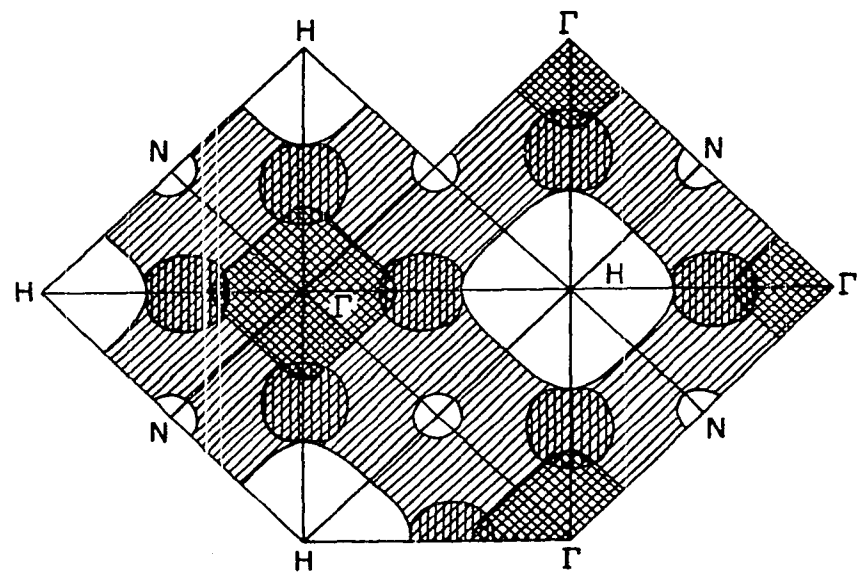


Figure 2. A $\langle 100 \rangle$ cross section for the Fermi surface of molybdenum as proposed by Lomer (3)

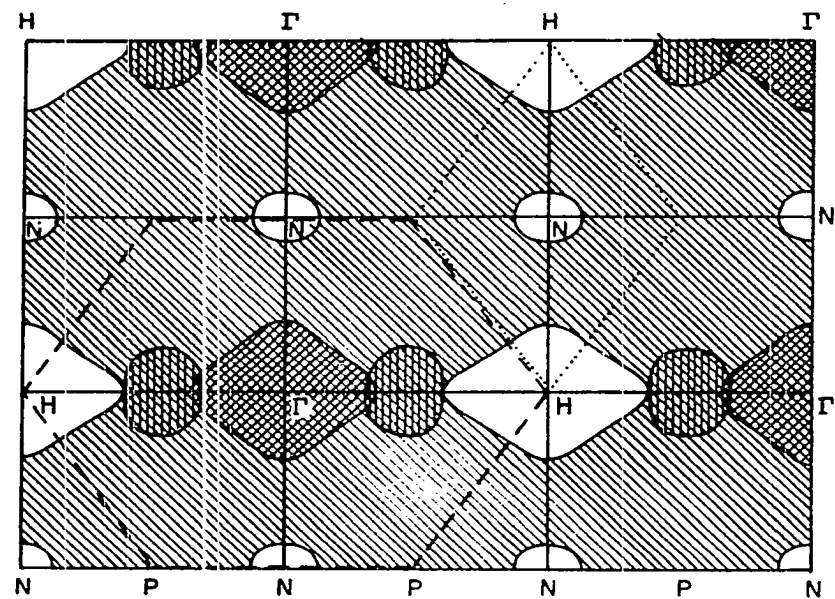


Figure 3. A $\langle 110 \rangle$ cross section for the Fermi surface of molybdenum as proposed by Lomer (3)

On the experimental side, the magnetoresistance measurements of Fawcett (7), of Fawcett and Reed (8), and of Alekseevskii, Egorov, Karstens, and Kazak (9) have shown that Mo is a compensated metal having equal numbers of electrons and holes. This means that the volume of the electron and the hole portions of the FS will be equal. Fawcett and Reed have shown there can be no more than about 10^{-4} open orbits per atom in Mo, which means that all sheets of the FS are simply-connected.

The anomalous skin effect measurements of Fawcett and Griffiths (10) were able to show that the FS area for a member of the Cr group transition metals is much less than the area of a sphere containing six electrons per atom. This suggests that the free electron model is a poor approximation to the Fermi surfaces of these metals.

Caliper dimensions along or near major symmetry directions were estimated by Jones and Rayne (11) and by Bezuglyi, Zhevago, and Denisenko (12) in their magnetoacoustic effect investigations. Herrmann (13) has obtained the effective masses of several orbits in Mo and W through the use of cyclotron resonance experiments. His results indicating a rather spherically shaped surface at H disagree with the hole octahedron found at H by other investigators.

The first de Haas-van Alphen (dHvA) frequency results obtained by using the torsion method with fields up to 18 kG on the Cr group metals were reported by Brandt and Rayne (14, 15). Their data for Mo came from orbits associated with the electron lenses. Sparlin and Marcus (16) presented dHvA torsion measurements on Mo and W in fields up to 34 kG. Their data included information about the electron lenses, the hole

ellipsoids, and partial information on the hole octahedra and the necks of the electron jack. No oscillations related to the body of the jack were observed. The pulsed-field dHvA results of Leaver and Myers (17) provide additional information about the jack and octahedron pieces, but their data exhibits a significant amount of scatter.

The recent radio-frequency size effect (RFSE) measurements of Boiko, Gasparov, and Gverdtseteli (18), of Cleveland (19), and Cleveland and Stanford (20) provide many caliper dimensions of the FS but the data presented in these two investigations differ by 6%, a discrepancy which is greater than the combined experimental errors claimed by the investigators. The RFSE data is most helpful in defining the shapes of the larger FS pieces, but is hard to obtain and interpret for the ellipsoids and lenses, where the RFSE signals lie close together in magnetic field value, and are difficult to separate and follow as a function of angle. The separation or gap between the jack and the octahedral FS pieces is estimated by Boiko et al. as 2.5% of the FH dimension, while Cleveland's estimate for the gap is 7%. The larger gap was explained by choosing a value of the spin-orbit parameter for Mo larger than that predicted by Mattheiss (21) for W. The effects of spin-orbit coupling should be more pronounced in W, where the gap is 5% (22), than in Mo.

This investigation of Mo was undertaken to provide complete and accurate data concerning the extremal areas associated with each of the pieces of the FS using the pulsed-field dHvA technique. The data will be used to obtain the FS radii and thus clarify the discrepancy existing in the current RFSE data, and to resolve the question about the size of the

gap between the j_g and the octahedron caused by the spin-orbit interaction.

II. EXPERIMENT

A. Impulsive Field Method

The de Haas-van Alphen effect is observed as oscillations in the magnetization of pure metal single-crystal specimens at liquid-He temperatures in strong magnetic fields. Onsager (23) worked out a simple theory giving the frequency of dHvA oscillations by applying the Bohr-Sommerfeld quantization rules to the motions of electrons in an applied magnetic field. The motion of electrons in k-space normal to the applied magnetic field is quantized into orbits which enclose areas proportional to the field. As the field is increased, these orbits cross the Fermi surface, depopulate, and therefore induce oscillations in the free energy and the magnetic moment of the system. The frequency F for these dHvA oscillations is

$$F = \frac{c\hbar}{2\pi e} A_O(E_F) = 104.7 \left(\frac{MG}{\text{\AA}^2} \right) A_O(E_F) \quad (1)$$

where $A_O(E_F)$ is an extremal cross-sectional area of the Fermi surface normal to the applied field. Experimental observations have shown that the dHvA oscillations are actually periodic in $1/B$ rather than $1/H$. Since the real space electron orbits have diameters extending over hundreds of lattice spacings, it is reasonable that the electrons should see an average field B inside the sample rather than the externally applied field H.

A detailed calculation of the amplitude dependence of the dHvA effect was made by Lifshitz and Kosevich (24), who applied statistical

mechanics to a system of independent electrons in an applied magnetic field. Dingle (25) has shown that if the effect of electron scattering due to collisions is included, the r^{th} harmonic term is multiplied by a factor K which is

$$K = \exp(-2\pi^2 r m^* k_B T_D / e \hbar H) \quad (2)$$

where T_D is an effective temperature, the Dingle temperature, related to the width of the Landau levels. This term is very important in drastically reducing the total signal strength of the dHvA effect in samples which are not pure or of good crystal perfection. Such is not the case with our samples.

The effect of conducting a dHvA experiment at a finite temperature is to cause thermal damping of the oscillations. This thermal damping factor I_r is given for the r^{th} harmonic by

$$I_r = \frac{X_r}{\sinh X_r} = 2X_r e^{-X_r} [1 + e^{-2X_r} + \dots]$$

where,

$$X_r = 2\pi r^2 \frac{k_B T}{\hbar \omega_c} = 2\pi^2 r \frac{m^* k_B T}{e \hbar H} \quad (3)$$

$X_r \approx 3r$ for free electrons in a field of 50 kG and at a temperature of 1 K. A recent article by Goid (26) includes this expression and provides a comprehensive review of the theoretical aspects of the dHvA effect and the additional experimental techniques possible.

In this study dHvA oscillations in Mo were observed using the impulsive field method developed by Shoenberg (27) in 1957. Since the basic

impulsive field apparatus was essentially the same as that used by Anderson (28), and Panousis (29), only a brief summary will be included here.

A single-crystal specimen of high-purity material was immersed in a dewar of liquid He, which was designed so that the sample was situated inside the center of a solenoid. This entire assembly was contained in a dewar of liquid N₂ for cooling. A block diagram of the experimental circuitry is shown in Figure 4. A bank of capacitors with 5040 mfd total capacitance was charged to a maximum of 2400 volts. Discharging these capacitors through the solenoid produced a peak field of about 180 kG with a pulse time of 21 ms. This time-varying magnetic field induced a voltage proportional to $dM/dt = (dM/dH)/(dH/dt)$ in a pickup coil surrounding the sample. The magnetic field values were determined from the voltage readings across a 0.01 Ω standard resistor. Both the field values and the pickup coil voltage were sometimes monitored through the use of a dual-beam oscilloscope and Polaroid pictures taken of this output for permanent pictorial data storage. The pickup coil voltage was also displayed on a Hughes Model 104 Memo-scope storage oscilloscope as a monitor of signal quality and strength for each pulse. Both the pickup coil voltage and the voltage from the standard resistor were appropriately amplified, converted from analog to digital signals and stored in a small memory for later readout onto punched paper tape.

The pickup coil consisted of 2000 turns of No. 50 insulated copper wire followed by $1\frac{1}{2}$ layers of cigarette paper and about 860 turns wound in a reverse manner. The final number of reverse turns was adjusted so

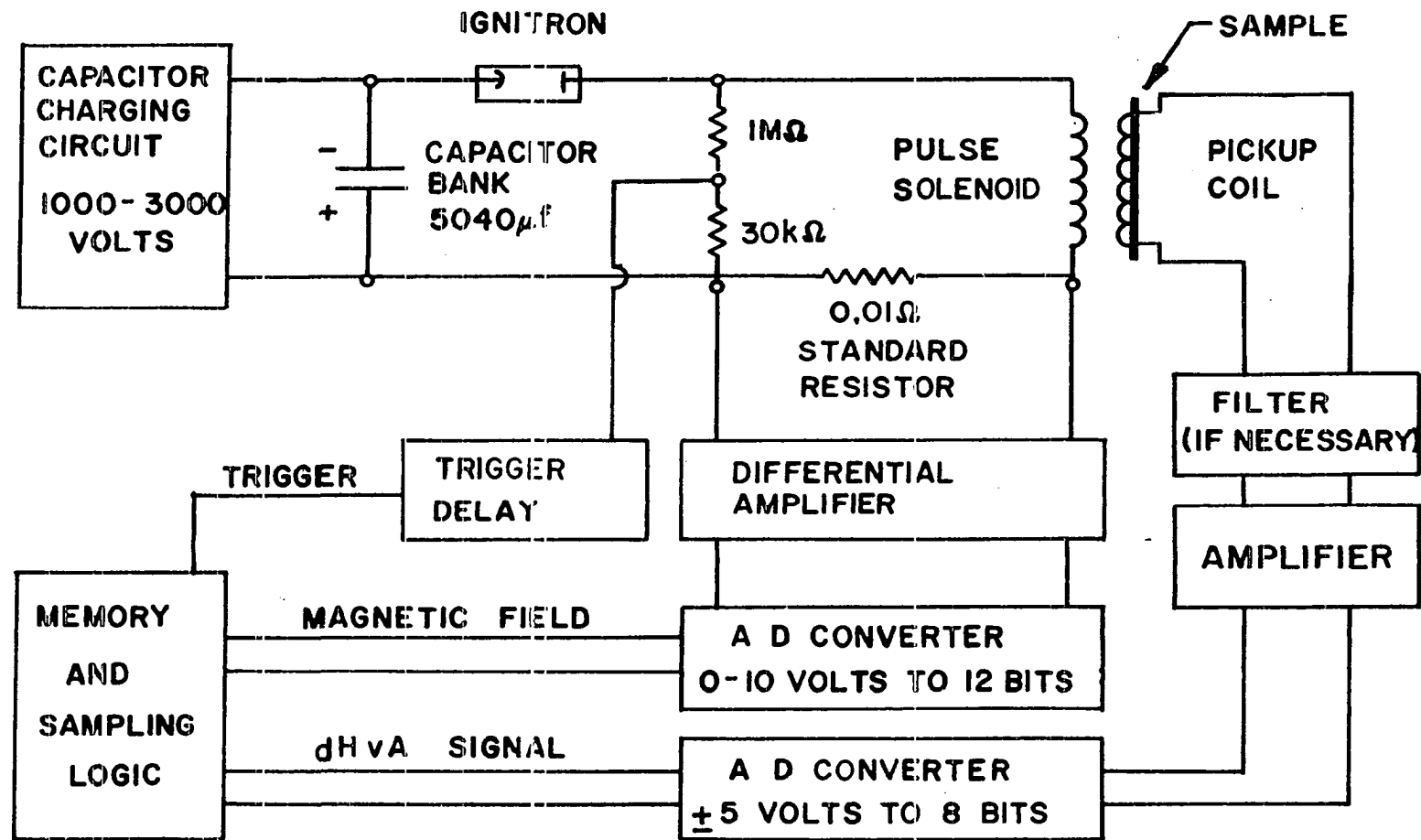


Figure 4. Block diagram of the experimental apparatus

that the "bucked" coil would not be able to detect the uniform field produced by the pulse solenoid, but would be able to pick up the changes in the magnetization of the sample residing inside. Since it is difficult to determine the exact number of turns necessary to completely "buck" the uniform time-varying field, a few turns of wire connected to an additional pair of leads from the top of the sample holder provided a signal which could be used to eliminate most of the remaining 24 Hz signal from the main field. These turns will be referred to later as the "bucking coil." Otherwise the dHvA signal appears as a ripple riding on an approximate one-half sine wave pulse coming from the solenoid.

B. Sample Preparation

The single-crystal samples used in this investigation were spark-cut from the same single-crystal rod of Mo from which Cleveland obtained samples for his RFSE experiment (19). The rod, purchased from Westinghouse Lamp Division, has a residual resistance ratio ($R_{300K}/R_{4.2K}$) of 5000 as determined by use of the eddy-current decay method (30). Two samples for this investigation were cut in the form of long, thin square bars -- one with a $\langle 100 \rangle$ crystallographic axis parallel to the length of the bar and the other with a $\langle 110 \rangle$ direction as the long axis. These samples were electro-polished to a diameter of 0.5 mm and a length of 5 mm, using a 6% solution of perchloric acid in methanol cooled with a mixture of dry ice and acetone.

A sample was mounted in a pickup coil housed in the drive wheel of the sample holder designed by Panousis and shown in Figure 5. The drive-wheel assembly was detached from the drive-shaft assembly by disconnecting

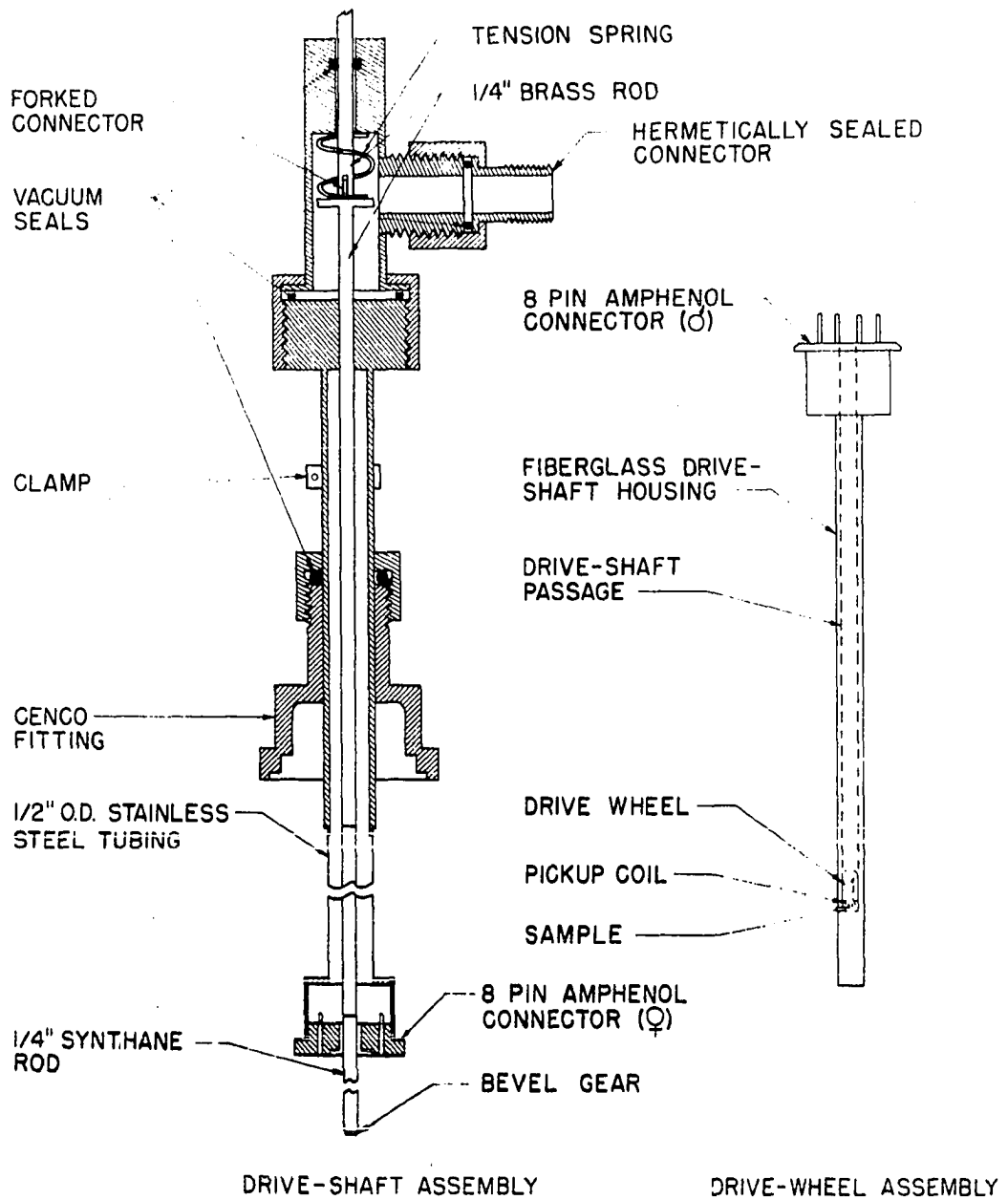


Figure 5. Rotating sample holder. The drive wheel assembly is fitted over the drive shaft and plugged into the Amphenol connector

the 8-pin Amphenol connectors and was mounted in a jig capable of keeping the plane of rotation of the drive wheel parallel to a Polaroid x-ray camera used for checking the orientation of the sample. The x-ray beam entered a small hole in the wheel parallel to its axle and struck the tip of the sample which slightly protruded out of the end of the pickup coil. The Laue back-scatter diffraction pattern obtained enabled us to determine the plane of rotation of the sample to within 1° of a chosen crystallographic direction.

C. Method of Obtaining Data

To obtain all of the frequency data possible from the sample, one would like to have some way of being able to discriminate against some frequencies while enhancing others. Then a higher amplitude signal from a given set of frequencies would enable a more accurate determination of the frequencies involved. Such discrimination is possible by exploiting the effects of the thermal damping term expressed in Equation 3. One data run for each orientation plane with a given sample was made at a temperature of 4.2 K using a peak field of 90 kG. This mode of operation enhanced the amplitudes of frequencies arising from the lenses and the ellipsoids. A second data run for each plane and sample was made at 1.2 K using a peak field of 170 kG. At this temperature and higher field the frequencies associated with the larger and/or higher effective mass orbits arising from the electron jack and the hole octahedron were brought out.

To contribute to the dHVA effect, electrons must complete one or several orbits in the real space of the crystal lattice. The size of the orbit is inversely related to the magnetic field strength. As the

magnetic field increases, more possible orbits begin having path lengths on the order of the mean free path λ , which is dictated by the sample purity as well as the temperature. Electrons which can complete an orbit contribute to the dHvA effect rather than being randomly scattered. The size of the orbit also depends upon the possible momentum states for the given electron as dictated by its position very near the FS in momentum space, and the direction of the applied magnetic field. Increased temperature causes the FS to become less sharply defined and increases the probability that the electron will be scattered onto a new path. The effective mass is a measure of the curvature of the particular orbit, and tells how easily the electron will be able to traverse the path that it is prescribed to follow. Temperature, effective mass and field strength all appear in the thermal damping term.

At 4.2 K and 90 kG peak field the total signal strength from the pickup coil was such that a gain of around 500 was needed to provide the ± 5 volts maximum signal to the analog-to-digital converter (ADC). A solid-state amplifier built by the Ames Laboratory Instrumentation Group especially for the dHvA investigation in lutetium by this author was satisfactory for this signal level. This amplifier, giving linear response from 1 kHz-100 kHz, provided a low-noise output except near its maximum gain of 50,000. Although the amplifier was non-linear at 24 Hz, the signal at this frequency due to the pulse solenoid was rejected by using the "bucking coil" mentioned in Section II-A. The "bucking coil" signal was reduced by a Dekatran transformer and fed to the first-stage differential amplifier in such a manner as to subtract out most of the

remaining pulse solenoid field which was detected along with the desired dHvA signal by the pickup coil, which was the other input to the same differential amplifier. Through the use of this complete amplifier we have provided a good wide-band dHvA signal of ± 5 volts maximum as input to the ADC without using any electronic or static filtering which could introduce noise.

At 1.2 K and the higher peak field the increased signal strength overloaded the stages ahead of the gain control on the amplifier mentioned above, so an older system was substituted. Here the pickup coil signal was the input to a Krohn-Hite Model 315AR electronic filter set with a band-pass of 5 kHz to 100 kHz. The filter was used to eliminate the 24 Hz pulse field signal and the associated "ringing" of the Infrared Model 603 amplifier which followed the filter in the circuit. This amplifier, which had a poorer frequency response than the newer one, was sufficient to provide a gain of around 80 with a reasonably good signal for the ADC.

For each plane of orientation of a given sample with respect to the magnetic field, and the temperature and peak field combinations mentioned above, data was obtained at 2° intervals as indicated by a counter dial connected by a set of gears to the $1/4''$ brass rod at the top of the sample holder shown in Figure 5. Backlash in the gears was eliminated by always advancing the counter to increasing angles from the beginning of a particular data run. Although the approximate symmetry direction of the sample for a certain dial reading was known and used to determine the angular interval over which data was taken, the final choice of the symmetry direction was made from the symmetry of the plotted frequency information

about some particular angle.

At each particular angular setting the solenoid was pulsed four times at 3 min. intervals for low peak field or 5 min. intervals for high peak field to allow the solenoid to reach thermal equilibrium. During two of these pulses the trigger delay was set so that the digital equipment would sample 8 msec of data occurring before the peak field ("rising field") and during the other two pulses for data occurring after the peak field ("falling field"). Each such pulse produced a set of 2008 values for the dHvA signal and 20 precise values for the magnetic field sampled in equal increments of time with a field value taken once every 100 dHvA signal samplings. The digital information was dumped from a small memory onto punched paper tape. This instrumentation is described in Reference 31.

Each data set was analyzed over a series of frequency ranges using the fast-Fourier transform program listed and discussed in Appendix B. A time constant τ of 40 μ sec was found appropriate to correct for the time delay introduced by the finite inductance of the "non-inductive" standard resistor, for eddy current effects, and for other effects, including electronics, capable of introducing a time delay. The true field H is related to the apparent (measured) field H' (see Reference 31) by

$$H = H' - \tau (dH/dt) \quad . \quad (4)$$

This correction produced dHvA results that were the same for both "rising field" and "falling field" data. Corresponding frequencies from each of the four data sets obtained at a given angle were averaged together to produce a single value which could be plotted on graphs of the data.

The number of degrees between symmetry points of the plotted raw data disagreed slightly with the dial readings of the sample holder. Since the $\frac{3}{4}$ " synthane rod of the sample holder (Figure 5) had a 14 tooth bevel gear which meshed into the 36 teeth of the drive wheel, the small number of teeth (which were cut as precisely as possible) could be a source of error in knowing the angle at which the data was taken. To correct for this error, the following procedure was adopted. The pulse solenoid was driven by a low-frequency oscillator run at maximum power. A lock-in detector and digital voltmeter measured the pickup coil voltage as a function of angle determined by the dial reading. The resulting cosine-like dependence of the voltage of the pickup coil in a uniform magnetic field was plotted along with a perfect cosine wave. The engaging of individual teeth as the angle was changed was visible on the plot. A table of actual angle for a given dial reading was constructed and these angle corrections applied to all the data. Such corrections are meaningful only if the shaft engaged the wheel at the same place for a given dial setting each time the sample holder was assembled, which was the case for this entire investigation.

Pulse solenoid G, used for all of the data presented here, was calibrated using the dHvA effect itself to make a dynamic calibration of the entire apparatus. The [111] frequency of the ν oscillation in tungsten, determined by O'Sullivan and Schirber (32) to be 98.8 ± 0.2 MG with the aid of NMR measurements of quasi-static fields, was used as the frequency standard. The field constant determined in this way was

$$K = 135.68 \pm 0.3 \text{ G/amp} .$$

This constant can be compared with one determined by Panousis (29) for the same solenoid, using lead as a standard, to be $K = 134.7 \pm 0.3$ G/amp with a systematic error of up to 0.5% because of sample misalignment, and with the static calibration of the entire apparatus made earlier by Phillips (33) which yielded a constant $K = 135.6$ G/amp. The static calibration and the dynamic calibration with tungsten agree within 0.1%. The 3/4% discrepancy quoted in Reference 22 seems to have been resolved by this additional calibration run.

The experimental data which is presented in the following section was subjected to the following criteria: (i) The frequency presented should be the average of corresponding frequencies found in the four data sets taken at that angle. (2) The frequency should be a member of a set of frequencies which are smoothly varying in value and amplitude as a function of angle of observation. (3) Sets of frequencies which can be identified as 2nd or 3rd harmonics, or sums and differences of other reported frequencies are omitted.

D. Frequency Results

The angular dependence of the fundamental dHvA frequencies corresponding to extremal cross-sectional areas of the Fermi surface of Mo are shown in Figures 6 through 11. In all of these figures the symbol (Δ) is used to represent data taken with the [110] sample while the symbol (o) is used for the [001] sample. The labels for the frequency branches are those used by Girvan et al. (22), except for the β used for lens frequencies, which do not exist in W.

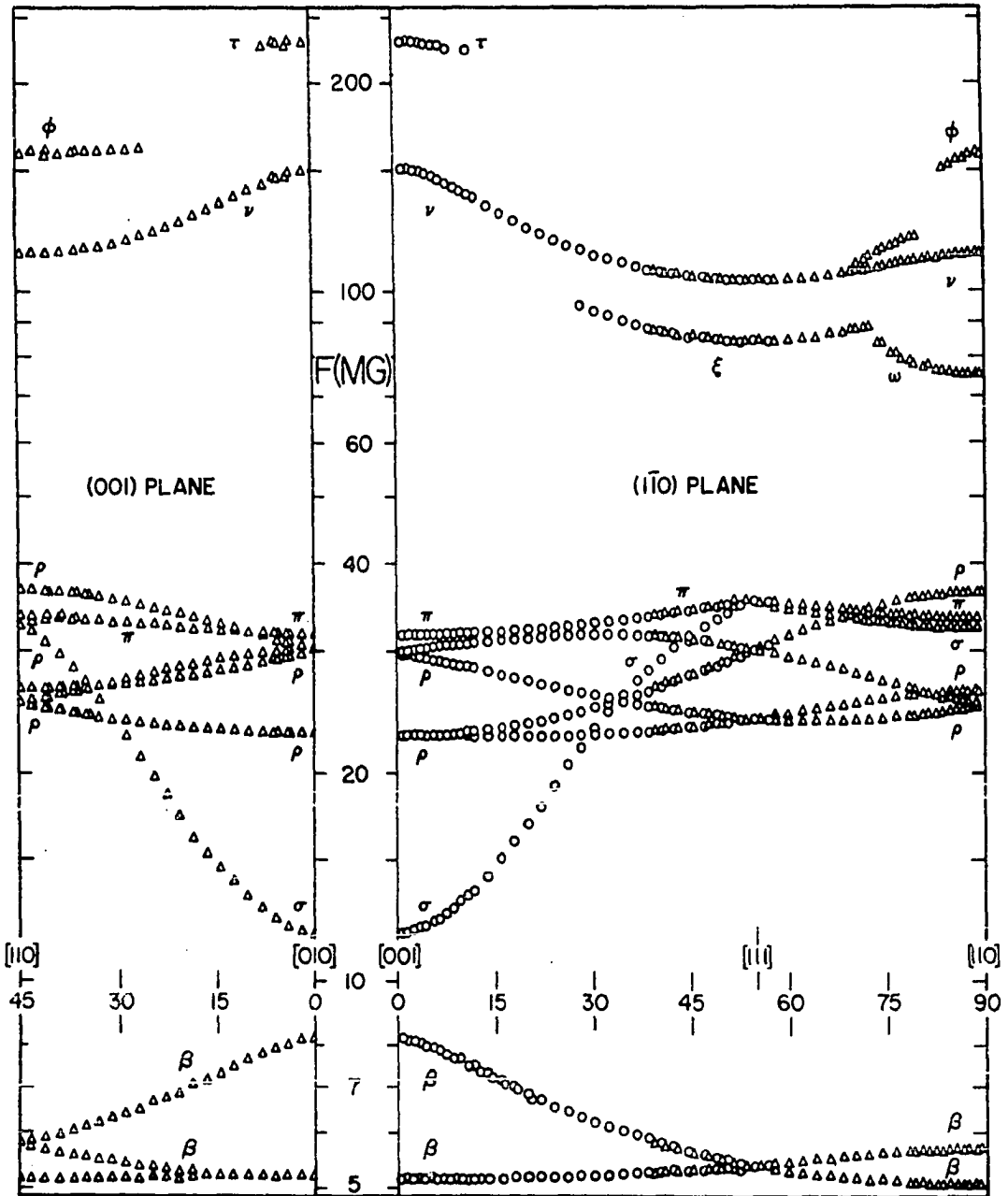


Figure 6. Variations of the fundamental de Haas-van Alphen frequencies in molybdenum for field directions in the (001) and (110) planes. The geometric axes for the rod shaped samples are: Δ , [110] and o, [001]

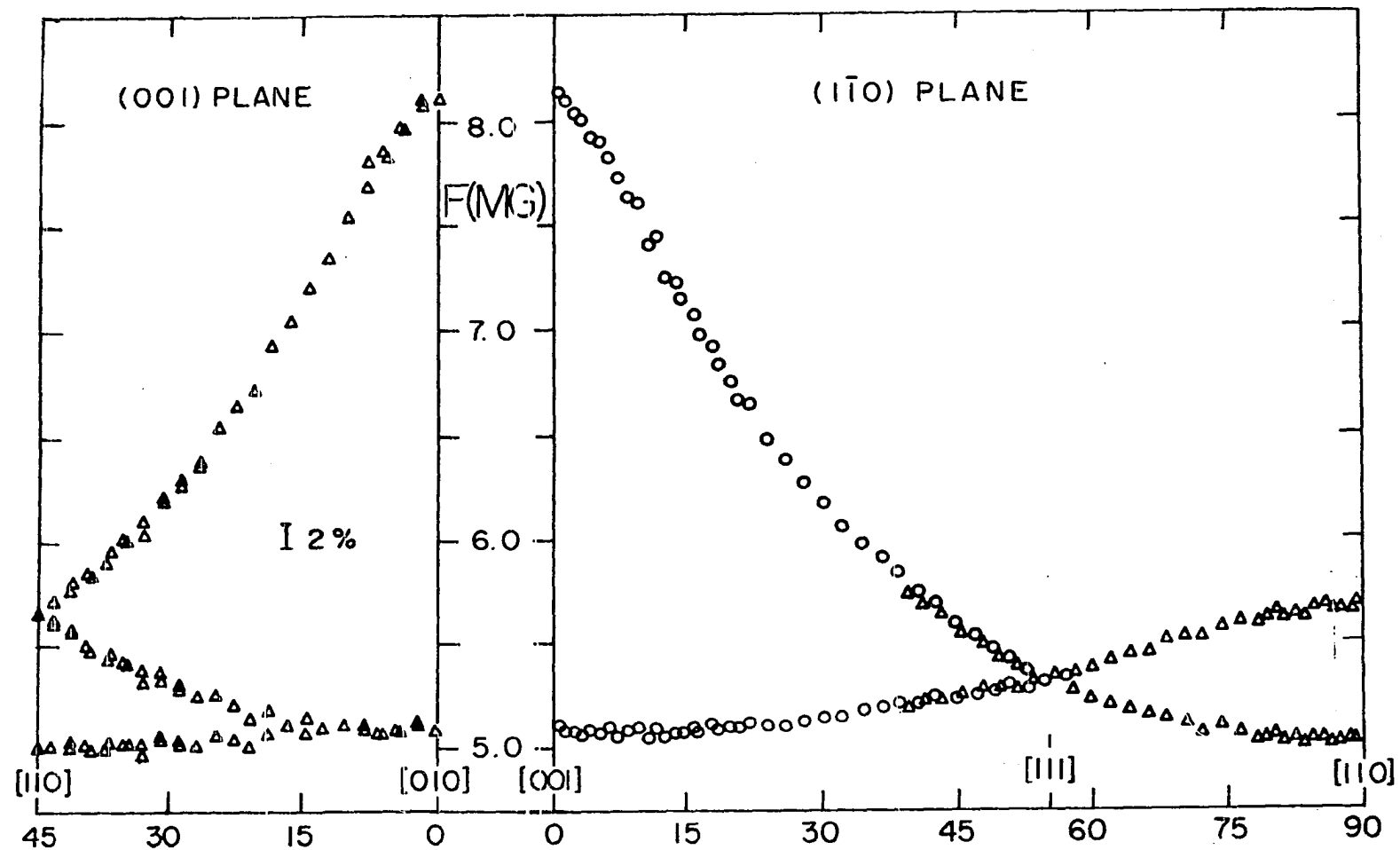


Figure 7. Linear plot of low frequency de Haas-van Alphen data for field directions in the (001) and (110) planes

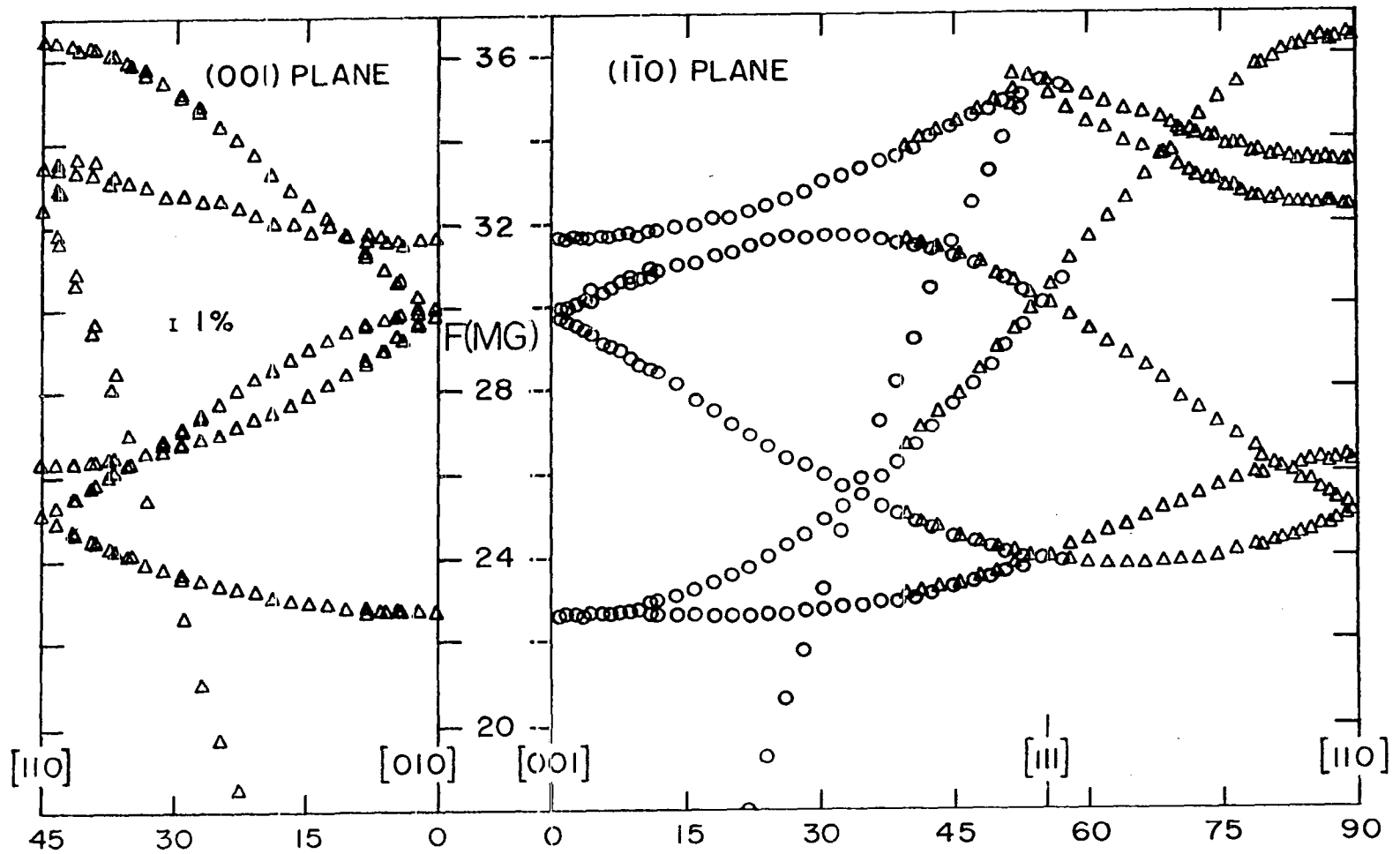


Figure 8. Linear plot of medium frequency de Haas-van Alphen data for field directions in the (001) and $(1\bar{1}0)$ planes

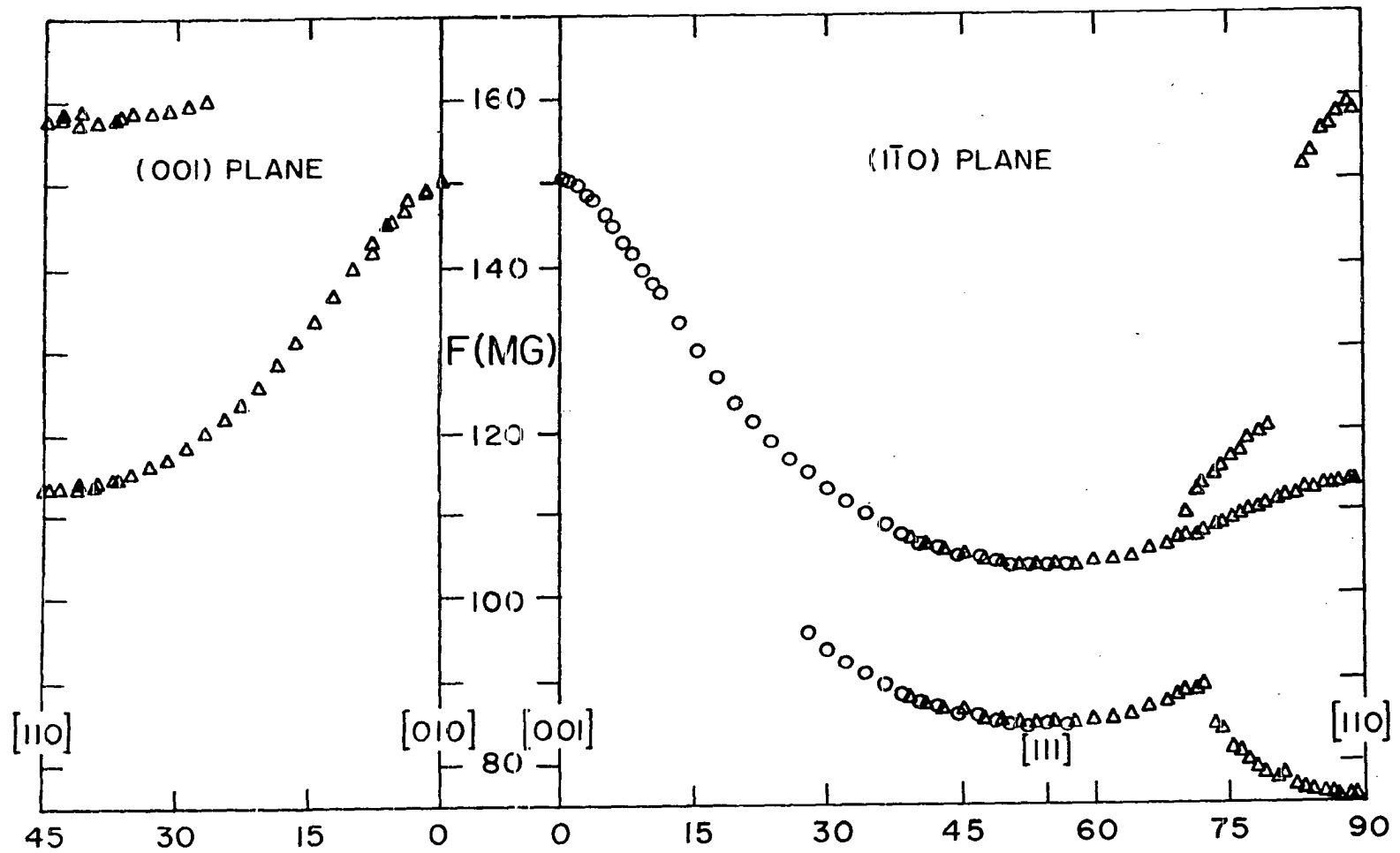


Figure 9. Linear plot of high frequency de Haas-van Alphen data for field directions in the (001) and (110) planes

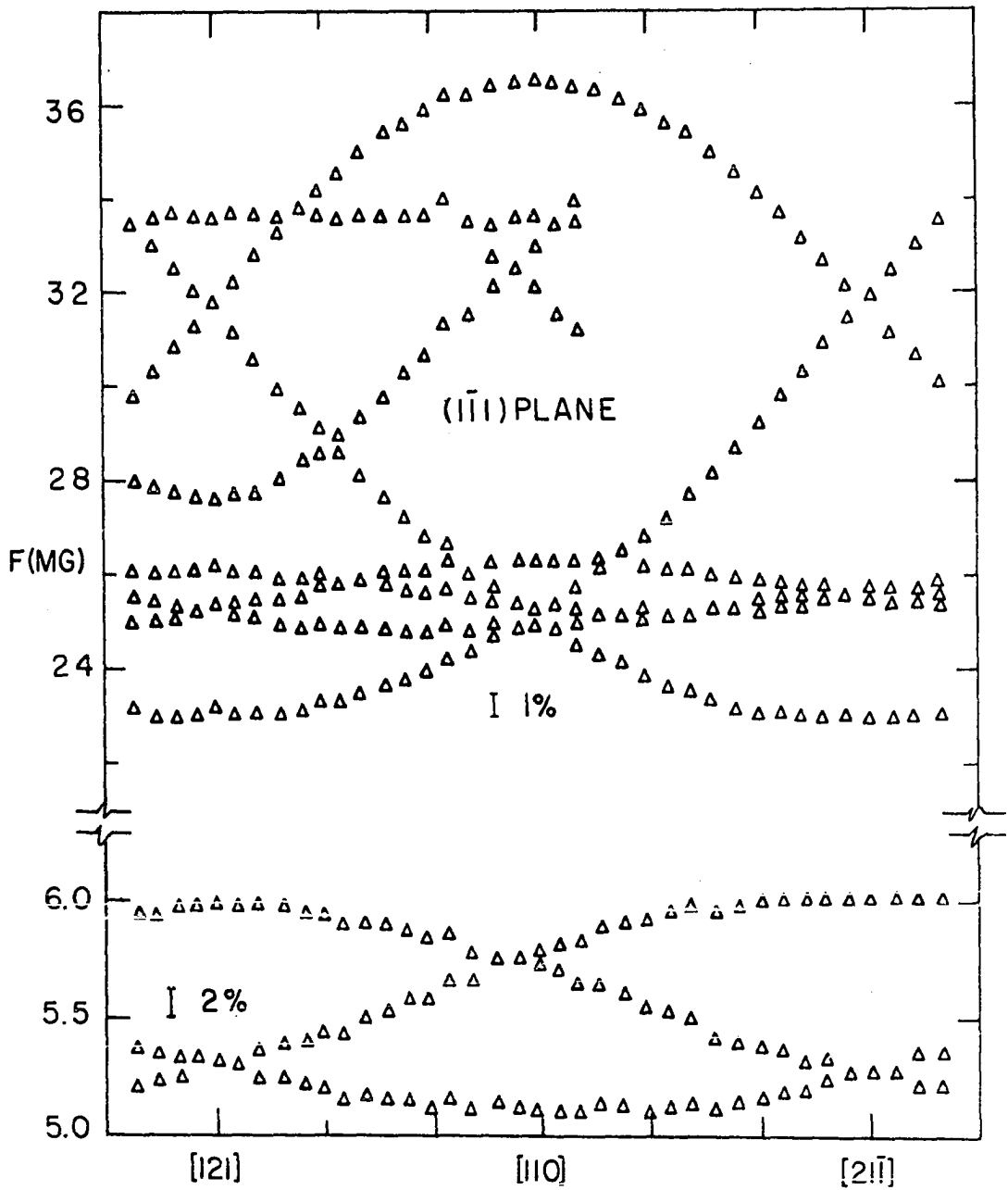


Figure 10. Linear plot of low and medium frequency de Haas-van Alphen data for field directions in the (111) plane. Note the break in scale for F (MG)

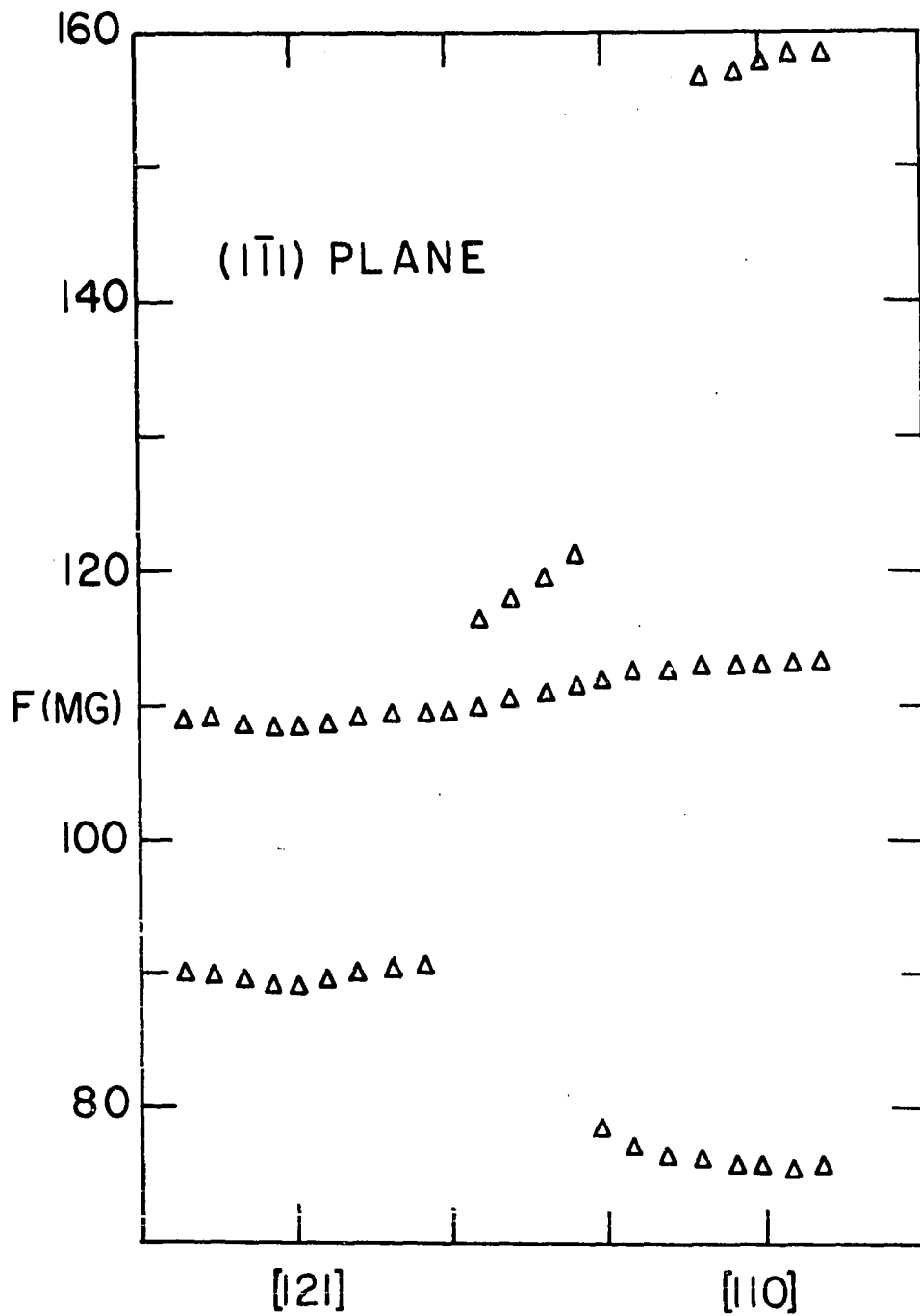


Figure 11. Linear plot of high frequency de Haas-van Alphen data for field directions in the $(1\bar{1}1)$ plane

The only correction applied to all of the data was that for angle calibration as described in Section II-C. In Figures 6 through 9 a symmetry direction was chosen for the corrected data taken over an approximately 60° range with a given sample and plane of orientation, and this data was folded about the chosen symmetry direction. Wherever the density of data as a function of angle is greater than one set of frequencies per approximate two degree interval, this folding process has taken place.

dHvA frequencies measured at a given symmetry direction should be identical, no matter which plane of orientation has been chosen or the sample being used. These frequencies, listed in Table I along with values determined by other experimentalists, can be compared for a given symmetry direction. Any discrepancies found are the result of misalignment of the sample in the plane of rotation, or the precision with which a given frequency can be determined (usually better than $\frac{1}{2}\%$, except for around 1% for low frequencies).

Figure 6 is a semi-logarithmic plot of all of the data obtained in the (001) and (1 $\bar{1}$ 0) planes. Figures 7, 8, and 9 are linear plots of most of the data appearing in Figure 6. The precision with which the samples were oriented can be seen from several aspects of this data. The folded data shows very little scatter as a function of angle near the [110] and [001] symmetry directions. The scatter in the folded data at [010] in the (001) plane is a function entirely of the choice of a symmetry direction at [110], since all of the information reported in the plane comes from the [110] sample. The scatter in the data appearing near [111]

Table 1. Experimental values for the dHvA frequencies (in megagauss) corresponding to extremal cross-sectional areas of the Fermi surface of Mo

| (001) PLANE | | | | (110) PLANE | | | | (111) PLANE | | | |
|-----------------|----------------------------|-------------------------|-------------------------|-----------------|----------------------------|-------------------------|-------------------------|-----------------|----------------------------|-------------------------|-------------------------|
| Label | Frequency this study | Other result (15) | Other result (16) | Label | Frequency this study | Other result (15) | Other result (16) | Label | Frequency this study | Other result (15) | Other result (16) |
| [010] direction | | | | [001] direction | | | | | | | |
| β | $5.19 \pm .03$ | 5.15 | 5.21 | β | $5.18 \pm .03$ | 5.17 | 5.33 | | | | |
| β | $8.21 \pm .04$ | 7.92 | 8.23 | β | $8.25 \pm .04$ | 8.00 | 8.53 | | | | |
| σ | $11.73 \pm .05$ | | 11.78 | σ | $11.73 \pm .05$ | | 12.30 | | | | |
| ρ | $22.7 \pm .1$ | | 23.0 | ρ | $22.7 \pm .1$ | | 23.0 | | | | |
| ρ | $29.9 \pm .1$ | | 31.2 | ρ | $29.9 \pm .1$ | | 31.1 | | | | |
| π | $31.65 \pm .1$ | | | π | $31.63 \pm .1$ | | | | | | |
| ν | $150.3 \pm .6$ | | | ν | $150.5 \pm .5$ | | | | | | |
| τ | 231 ± 3 | | | τ | 232 ± 3 | | | | | | |
| [110] direction | | | | [110] direction | | | | [110] direction | | | |
| β | $5.12 \pm .02$ | | 5.08 | β | $5.12 \pm .02$ | 5.03 | 4.95 | β | $5.10 \pm .1$ | 5.04 | |
| β | $5.75 \pm .02$ | 5.70 | 5.76 | β | $5.76 \pm .03$ | 5.63 | 5.65 | β | $5.74 \pm .1$ | 5.75 | |
| ρ | $25.1 \pm .1$ | | 26.18 | ρ | $25.1 \pm .1$ | 24.2 | 25.76 | ρ | $25.0 \pm .2$ | > 25.8 | |
| ρ | $26.3 \pm .1$ | | 27.96 | ρ | $26.3 \pm .1$ | | 29.5 | ρ | $26.3 \pm .2$ | | |
| σ | $32.4 \pm .1$ | | 32.25 | σ | $32.35 \pm .1$ | | 30.8 | σ | $32.4 \pm .2$ | | |
| π | $33.4 \pm .15$ | | 34.0 | π | $33.4 \pm .1$ | | | π | $33.5 \pm .2$ | | |
| ρ | $36.4 \pm .1$ | | 40.1 | ρ | $36.45 \pm .1$ | | 38.5 | ρ | $36.4 \pm .2$ | | |
| ω | $76.0 \pm .4$ | | | ω | $76.0 \pm .4$ | | | ω | $75.6 \pm .4$ | | |
| ν | $113.2 \pm .4$ | | | ν | $113.4 \pm .4$ | | | ν | $113.2 \pm .4$ | | |
| ϕ | 158.0 ± 1.0 | | | ϕ | 158.7 ± 1.0 | | | ϕ | 158 ± 1.5 | | |

[112] direction

| | |
|----------|----------------|
| β | $5.27 \pm .03$ |
| β | $6.04 \pm .04$ |
| ρ | $22.9 \pm .1$ |
| ρ | $25.4 \pm .1$ |
| ρ | $25.7 \pm .1$ |
| σ | $26.3 \pm .3$ |
| ρ | $31.7 \pm .1$ |
| π | $33.3 \pm .1$ |
| ξ | $90.5 \pm .4$ |
| ν | $109.5 \pm .4$ |

[121] direction

| | | |
|----------|-----------------|------|
| β | $5.28 \pm .1$ | 5.28 |
| β | $5.98 \pm .1$ | 5.97 |
| ρ | $23.0 \pm .15$ | |
| ρ | $25.2 \pm .3$ | |
| ρ | $26.1 \pm .4$ | |
| σ | $27.6 \pm .4$ | |
| ρ | $31.7 \pm .2$ | |
| π | $33.6 \pm .3$ | |
| ξ | 89.0 ± 1.0 | |
| ν | 108.2 ± 1.0 | |

[111] direction

| | | | |
|----------|----------------|------|-------|
| β | $5.41 \pm .03$ | 5.37 | 5.45 |
| ρ | $24.0 \pm .1$ | | 24.45 |
| ρ | $30.2 \pm .1$ | | 30.9 |
| σ | $35.2 \pm .2$ | | 36.4 |
| π | $35.4 \pm .2$ | | |
| ξ | $84.7 \pm .4$ | | |
| ν | $103.5 \pm .4$ | | 108.6 |

in the $(1\bar{1}0)$ plane, where the data has come from both samples, is a measure of our ability (1) to place each of the two samples in the $(1\bar{1}0)$ plane, (2) to apply the same angle correction for a given dial reading on the sample holder, and (3) to pick a symmetry direction for the folded data at 0° and 90° . Once these three things have been done, there are no additional variables at our disposal to enable us to do a better job of making the data from these two samples overlap.

In Figure 6 the frequency branches labelled β arise from orbits around the 6 electron lenses situated at each of the symmetry points Δ in the Brillouin zone. The ρ branches arise from the orbits associated with the 12 ellipsoids situated at the symmetry points N. The single frequency branch ν comes from orbits around the 6 equivalent octahedra located at the points H.

The rest of the orbits can be identified with the electron jack at Γ in the center of the zone. The frequency branches π and σ correspond to orbits around the balls and necks respectively. The τ frequency branch arises from a central orbit around four balls of the jack. The ω frequency branch, due to a non-central orbit containing part of two adjacent balls and the body between them, is shown in the $(1\bar{1}0)$ plane in Figures 6 and 9 and should also appear in the (001) plane at $[110]$ and exhibit some angular dependence. Although this frequency branch probably existed in the raw data, it could not be sorted out from the pattern of 2nd and 3rd harmonics arising from data in the 25-40 MG range, and was therefore omitted. The two data points between 71° and 73° which look like they could be part of the ξ frequency branch actually belong to the end of the

ω frequency branch, as determined by comparison of their amplitudes with those of nearby data points in the two branches. Then the unlabelled frequency branch which appears at angles between the ξ and ϕ frequency branches can be explained as arising from a non-central orbit probably encompassing the body and one ball of the jack as an intermediate between the ξ frequencies coming from orbits around the body of the jack, and the ϕ frequencies coming from an orbit encompassing two of the balls and the body between them.

Data for the $(1\bar{1}1)$ plane is shown in linear plots in Figures 10 and 11. Note the break in the frequency scale in Figure 10. The right half of this figure is missing some data because data was taken over a shorter angular range at high field for this plane. The angular scale shown with tick marks every 10° was that used in determining a symmetry direction for three frequency branches chosen for use in inversion of ellipsoid and octahedron data (see Sections III-B and III-C). It is obvious that this data does not have the symmetry which it should, and therefore, the data has not been folded about the symmetry direction. However, this plane is extremely sensitive to small misalignments of the sample. If the sample is only 1° out of the $(1\bar{1}1)$ plane, the four frequencies which should all cross at 25.1 MG at exactly $[110]$ will not cross there, but will be shifted. The new crossing point for those frequency branches lying entirely between 24 and 27 MG in this plane can be greater than 10° from $[110]$. These three frequency branches in the data exhibit the most deviation from the expected symmetry pattern. Notice also the 2° uncertainty in the exact position for the $[110]$ direction. A second attempt to

orient the sample in this plane resulted in data almost identical to that presented here. We estimate that an alignment error of less than 1° can account for the symmetry problems shown by this plane, and use this fact along with the success of our data in the other two orientation planes to determine the uncertainty of the plane of orientation of a sample with respect to the magnetic field as 1° or less.

III. DISCUSSION OF THE RESULTS

A. General Considerations

The experimental information about the FS of a metal obtained through the detection of the dHvA frequencies is in terms of extremal cross-sectional areas of the FS normal to the applied magnetic field (Equation 1). The FS information that one would really like to see would be the specification of the radius vector to any point that lies on the surface. Such information would constitute a complete geometric description of the Fermi surface.

A quick qualitative comparison of the Fermi surfaces of W and Mo can be made through the use of Figure 12. The plot exhibits the Mo data from Figure 6 as solid lines and the W data of Girvan et al. (22) as dashed lines. The band structures for these metals are similar, so qualitatively similar Fermi surfaces are to be expected, except that W does not have lenses. The octahedra of both metals compare rather closely in size. The electron jack of Mo is quite a bit larger in body, ball and neck dimensions than in W. The lenses in Mo add a small contribution to the volume of the total electron FS pieces. The hole ellipsoids are much larger in Mo than in W, which is to be expected if the hole volume is to exactly compensate for the increased electron jack volume and the volume due to the lenses. Remember that for each of these compensated metals the volumes of the electron FS pieces must equal the volumes of the hole FS pieces.

There are at least three ways of relating FS dimensions to the dHvA data. (1) An accurate band structure calculation for Mo could be done from which FS radii can be obtained. The intersection of a plane with

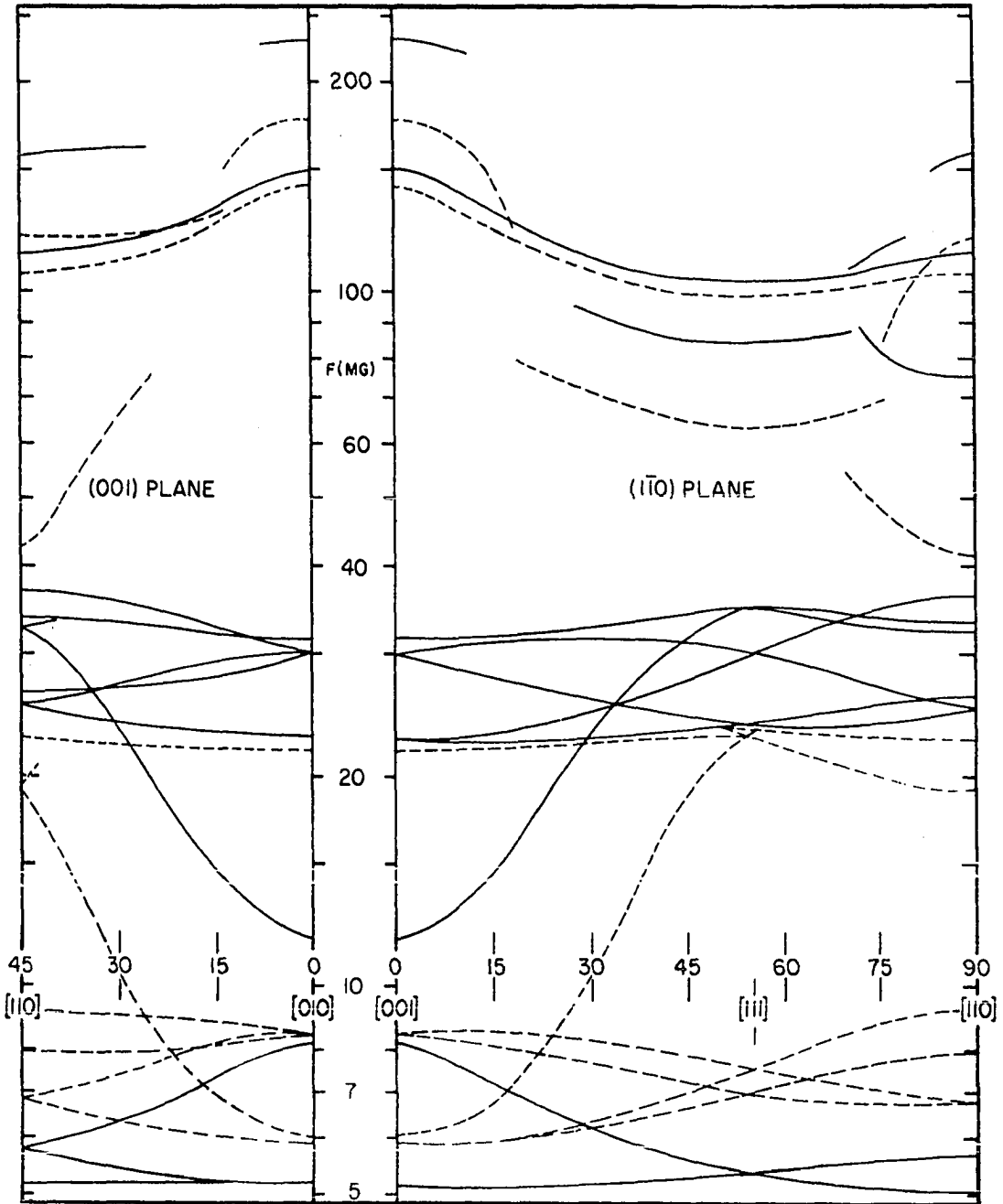


Figure 12. Comparison of the de Haas-van Alphen frequencies of molybdenum and tungsten. The solid lines indicate molybdenum data from this study. The dashed lines indicate tungsten data from Reference 22

these radii can be integrated to determine the extremal areas one would expect to see in a dHvA experiment. The calculated areas can be compared with the experimental data and conclusions drawn about the accuracy of the band structure calculation and about how consistent the data really is.

(2) The FS areas could be converted into FS radii by some inversion technique. The hole ellipsoids at N and the hole octahedron at H satisfy the requirements needed to invert the dHvA data to obtain FS radii by using computer programs (34) implementing the Mueller inversion scheme (35).

These requirements for a unique inversion are that the surface must:

(a) be closed, (b) have a center of inversion symmetry, and (c) have a unique radius vector from that center. The lenses and jack fail requirements (b) and (c) respectively. (3) Simple or complex geometrical models can be devised at the appropriate symmetry point and used to predict experimental extremal cross-sectional areas. Sometimes the models can be revised until the desired accuracy of fit to the data is obtained.

An accurate band structure calculation has not yet been done nor has anyone predicted the dHvA frequencies as a function of magnetic field direction, except for symmetry directions. The APW calculations of Loucks (4) show qualitative agreement with the Lomer model for Mo, but the numerical results are questionable (see Section 1), so we will not compare our results with theoretical numbers.

The inversion scheme programs (34) will be used for the octahedra and the ellipsoids, but can not be used on the lenses and jack. The difficulties encountered in applying the inversion scheme and the criteria for picking the number of coefficients to use in the expansion will be

discussed shortly.

The 11 parameters of the fairly complex geometrical model for the W jack (see Reference 22) were adjusted in an attempt to fit the Mo data. A program used to generate this model was obtained from Dr. R. F. Girvan (36). While a few of the parameters could be successfully chosen to fit data coming from the body of the jack, no complete set was found that would even come within 5% of all the data points available at high symmetry directions. The choosing of the parameters is done by hand, since the equation is highly non-linear and the data are compared to integrals of the equation. This is not to say that a better set of 11 parameters does not exist, but the likelihood of finding this better set is quite small. Therefore, we shall attempt to use only very simple geometrical models to determine the dimensions of the jack and compare these with the results of other experiments.

A significant error exists in the Appendix B of Reference 34 containing the inversion programs. The Kubic Harmonic Expansion Coefficients, necessary for use with the O_h group inversion programs, has the leading two columns of each page missing. These columns should contain "sign" information about several of the numbers in the table. The first coefficient in the table also has the digit "1" dropped, which is necessary to obtain any results at all from programs using this information. A partial list of these coefficients in Reference 37 was used in detecting this error and the complete list of correct values was later obtained from the program authors. One parameter, the cone angle, must be set at or near 90.0° rather than at 0° , which one might try as a first guess, since no "usual"

value is quoted to guide the user of the programs.

The symmetry at N is such that eight frequency branches for the ellipsoids are obtained from data taken in only the (001) and ($1\bar{1}0$) planes. The number of expansion coefficients uniquely determined by the data supplied to the inversion program is related not only to the number of data points, but also to the angular position at which the points are taken. Choosing to take data only in planes of high symmetry does place a limit on the number of coefficients which should be used for inversion. A general feeling for this limit develops as attempts are made to determine increasingly larger number of coefficients. For the octahedron at H, the full cubic symmetry O_h applied. Here data taken in only the (001) and ($1\bar{1}0$) planes can not yield unique coefficients whose l value is greater than 18 (11-terms) (see Reference 37). Including additional data from the ($1\bar{1}1$) plane should increase this number of terms somewhat. The number of coefficients used in the final fit for each surface was determined by the lesser of the maximum number obtainable and the point at which addition of more coefficients would yield no further reduction in the standard deviation of the data from that predicted by the inversion scheme.

B. Hole Octahedron at H

The data relating to the octahedron, labelled v in Figure 6, from all three planes was used with the inversion scheme for the O_h symmetry group to obtain the FS radii for this piece of the total surface. The symmetry directions chosen for the ($1\bar{1}1$) plane are those which are shown in Figure 10.

A plot of the normalized frequency difference (in percent)

$$\Delta F = 100(F_{\text{inversion}} - F_{\text{data}})/F_{\text{data}} \quad (5)$$

as a function of angle and ℓ value is shown in Figure 13. The frequencies from the dHvA data were interpolated to integer angles for comparison with the integer angle output of inversion frequencies. The inversion scheme fit all of the data within 0.8% and most of the structure in this curve can probably be attributed to the scatter in the input dHvA frequencies, rather than to the errors in fitting by the inversion program.

As indicated in Section III-A, $\ell = 18$ with 11 coefficients is the highest unique fit for data from only the (001) and (1 $\bar{1}$ 0) planes. Using the radii determined from an $\ell = 18$ fit using data from all three planes as a standard, we compare the radii obtained for higher values of ℓ . A plot of these radii differences (in percent)

$$\Delta R = 100(R_{\ell = 20, 22, \text{ or } 24} - R_{\ell = 18})/R_{\ell = 18} \quad (6)$$

as a function of angle is shown in Figure 14. We notice that none of the radii have changed by more than $\pm 2\%$ as the ℓ value has been increased. From the number of peaks and valleys in the $\ell = 24$ curve, we gather that the fit is probably sensitive to large changes in radii (surfaces of high curvature) only for changes occurring over angular intervals of approximately 10° or more.

The cross sections of the octahedron pieces of the FS for the (001), (1 $\bar{1}$ 0), (1 $\bar{1}$ 1) and (1 $\bar{1}$ 2) planes are shown in Figures 15 and 16 for the $\ell = 24$ fit involving 19 coefficients. The 19 coefficients are listed in Appendix A. The $\ell = 24$ fit was chosen as the cut-off point where using

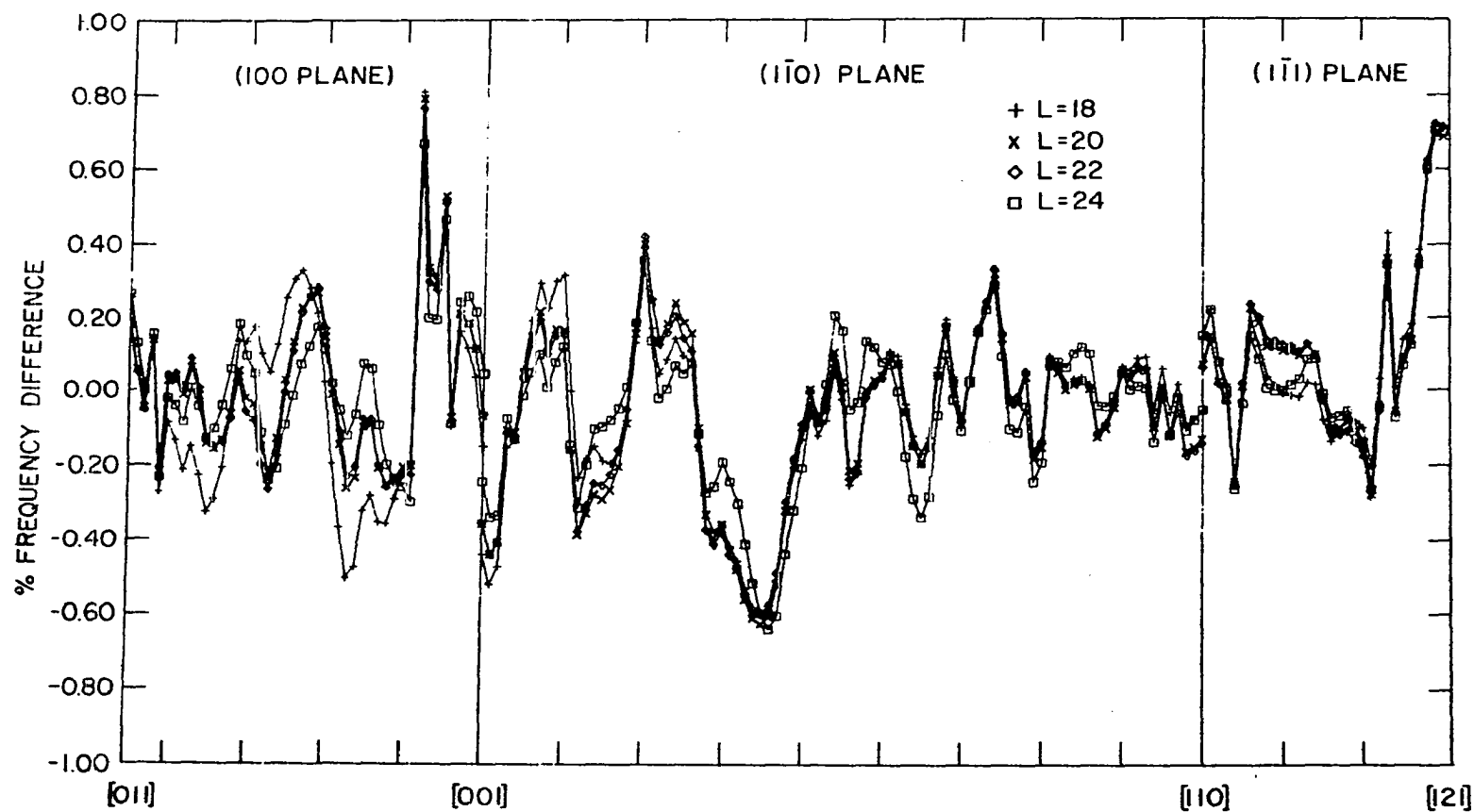


Figure 13. Plots of the normalized frequency difference (in percent) $\Delta F = 100 (F_{\text{inversion}} - F_{\text{data}}) / F_{\text{data}}$ as a function of angle and l value

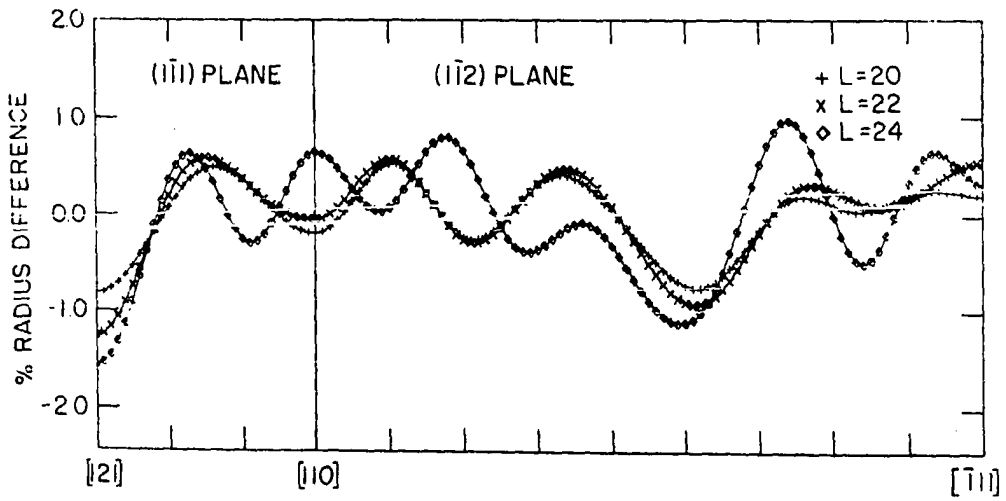
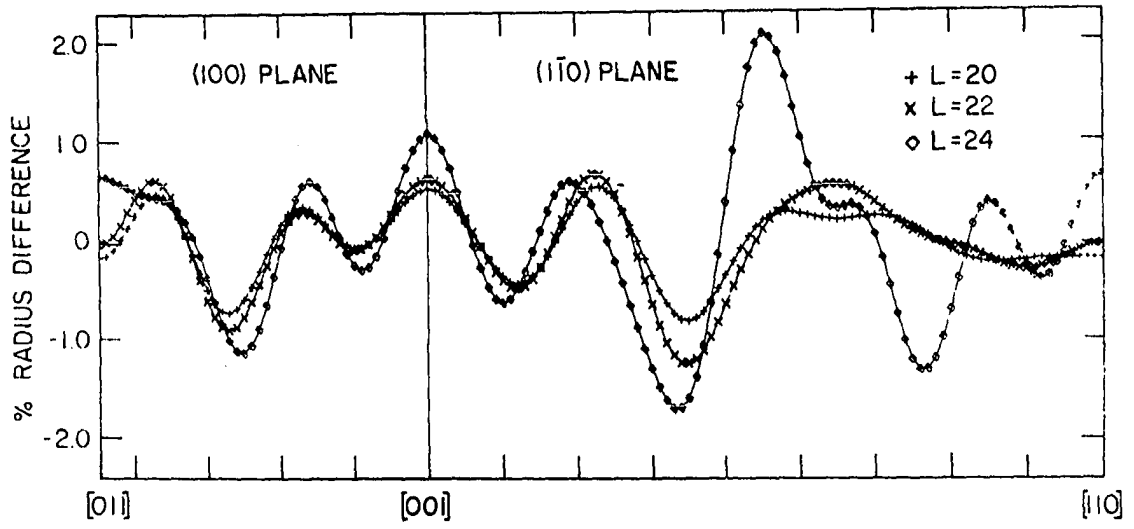


Figure 14. Plots of the radii differences (in percent)
 $\Delta R = 100 (R_L = 20, 22 \text{ or } 24 - R_L = 18) / R_L = 18$ as a function
 of angle

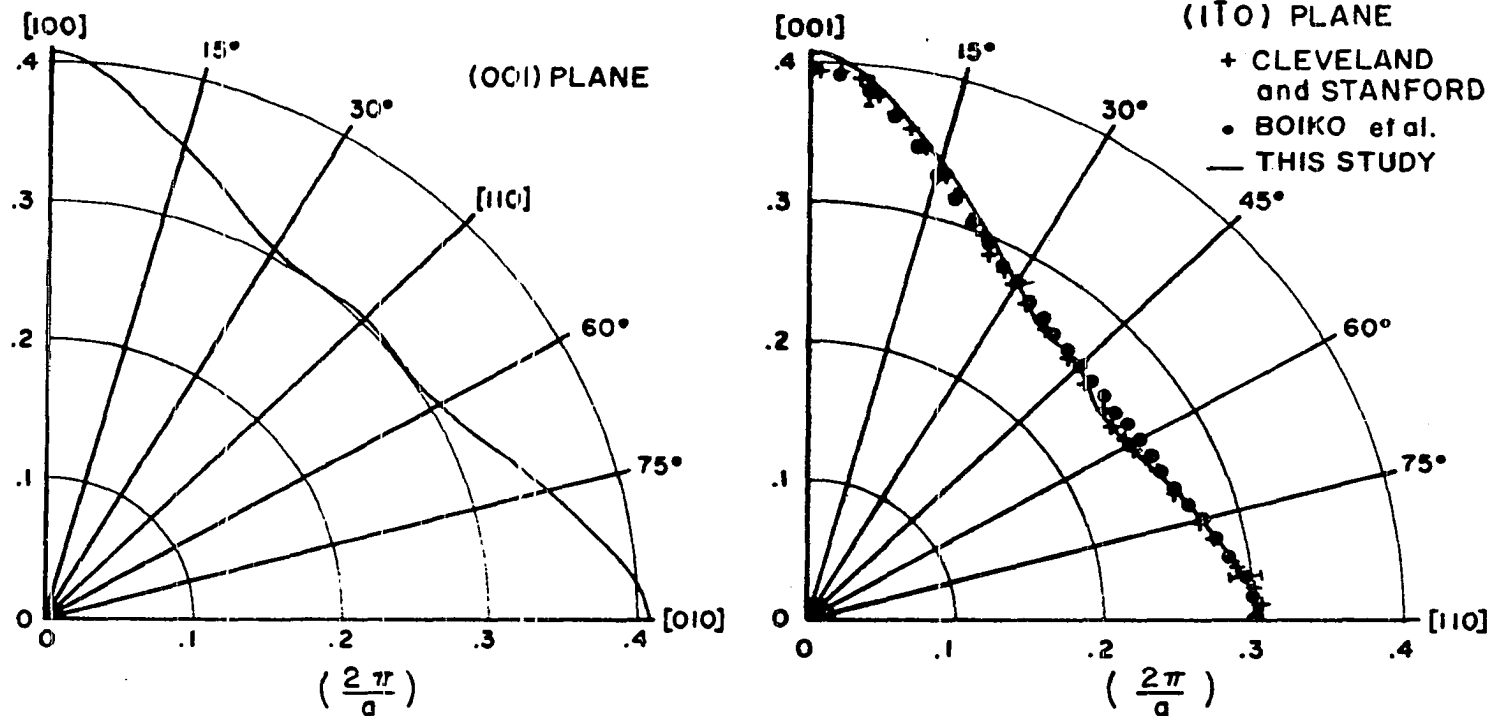


Figure 15. Cross sections of the hole octahedron Fermi surface at H for the (001) and (110) planes. The RFSE data of Boiko et al. (18) has been plotted directly, while the RFSE data of Cleveland and Stanford (20) has been multiplied by 1.05 before plotting

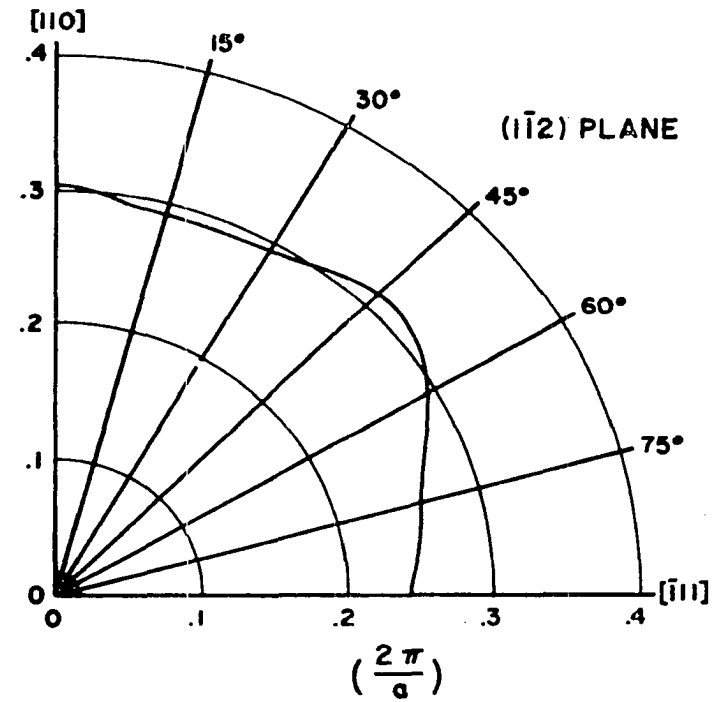
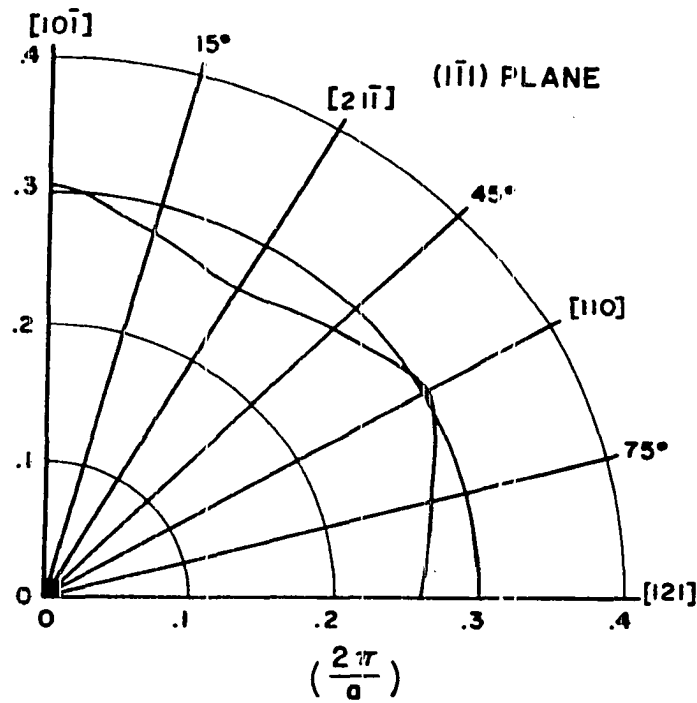


Figure 16. Cross sections of the hole octahedron Fermi surface at H for the $(\bar{1}\bar{1}1)$ and $(\bar{1}\bar{1}2)$ planes

additional coefficients would not further reduce the standard deviation of the input data from that predicted by the inversion scheme. In the $(1\bar{1}0)$ plane in Figure 15, the RFSE caliper dimensions of Boiko et al. (18) have been plotted directly, while the dimensions obtained by Cleveland (19) and Cleveland and Stanford (20) have been multiplied by 1.05 before plotting. In the $(1\bar{1}0)$ plane the RFSE calipers for this piece of the FS are also the FS radii, except for angles near 50° , where the calipers would not be able to measure the slight dimple in the surface near these angles, and would, therefore, yield larger values than those of a plot of radii. At the $[001]$ direction the RFSE measurement is probably more accurate than the radii indicated by the inversion scheme, since the surface has high curvature near this direction.

Radii obtained from inversion of the dHVA data for the octahedron are tabulated for major symmetry directions in Table 2. The volume

Table 2. Radii for the hole octahedra in units of $(2\pi/a)$

| Direction | Inversion scheme radii | Others | |
|-----------------------|------------------------------|-----------------------------|---------------------------------|
| | | Boiko <u>et al.</u> (18) | Cleveland & Stanford (20) |
| $\langle 100 \rangle$ | .407 | .395 | .376 |
| $\langle 110 \rangle$ | .304 | .30 | .290 |
| $\langle 111 \rangle$ | .244 | .255 | .238 |
| $\langle 112 \rangle$ | .262 | .27 | .253 |

calculated using the inversion coefficients in a suitable program from Reference 34 was found to be 0.90 \AA^{-3} . This compares with 0.90 \AA^{-3} obtained by Boiko et al. (18) and 0.95 \AA^{-3} estimated by Sparlin and Marcus (16).

From this comparison with RFSE data we would agree with the measurements of octahedron radii by Boiko et al. to within errors of the experiments. Our plots of radii are probably accurate to $\pm 1\%$, except near the [001] direction, where the errors are probably closer to $\pm 2\%$.

The RFSE octahedron dimensions reported by Cleveland and Stanford (20) for the (110) plane were integrated by computer and yielded an area of $0.98 \pm 0.02 \text{ \AA}^{-2}$. When these dimensions were increased by 1.05, the resulting area matched the $1.083 \pm .01 \text{ \AA}^{-2}$ cross-sectional area for this direction obtained in the present experiment. From the manner in which these RFSE dimensions, plotted in Figure 15 after being multiplied by 1.05, compare with this study and with the RFSE calipers of Boiko et al., it would appear that the work of Cleveland and Stanford was subject to a systematic error of 5%. In Reference 20 they are aware that their results do not agree with those of other investigators and have considered in detail the sources of error in their experiment. Their investigation was thorough and does not contain obvious weaknesses. Using samples cut from the same single-crystal rod, we present experimental data which disagrees with their work, but must conclude as they did concerning their work that "the reasons for the disagreement are at present unknown." (20).

C. Hole Ellipsoids at N

All of the data labelled ρ in Figure 6 comes from the hole ellipsoids at N. These frequency branches, four from the (001) plane and four from

the $(1\bar{1}0)$ plane, were used as input to the inversion scheme for the $V_h (D_{2h})$ symmetry group along with the two frequency branches from the $(1\bar{1}1)$ plane which are the most insensitive to small errors in sample orientation. These latter two were the branches connecting the 31.7 MG frequency at the $\langle 112 \rangle$ directions with either the 36.4 MG or 25.0 MG frequencies at the $[110]$ direction as shown in Figure 10 with the choice of symmetry directions as indicated on that figure.

There are 12 symmetry points N in the Brillouin zone in Figure 1, and generally six frequency contributions to the dHvA effect for an arbitrary magnetic field direction from a FS piece appropriately situated at equivalent pairs of these points, but we must concentrate on a single point N in order to obtain results from the inversion scheme. The problem can be viewed as one in which we place an ellipse with axes a, b, and c along three coordinate directions k_x , k_y , and k_z with point N at (0,0,0) and act as if the four frequency branches from the (001) plane had actually come from four separate planes of observation on the single ellipse, and similarly for any other experimental plane of observation. As a check on the data assignments made, computed dHvA data generated from a perfect ellipsoid were fed into the inversion scheme and checked with the areas calculated by the inversion scheme. The appropriate choice of spherical mapping coefficients could also be determined with this ideal data.

The three principal cross sections for the ellipsoid pieces situated at symmetry points N are shown in Figure 17 with different plotting symbols for each of the cross sections. The solid curves are plots of perfect ellipses having major and minor axes identical to those obtained from the

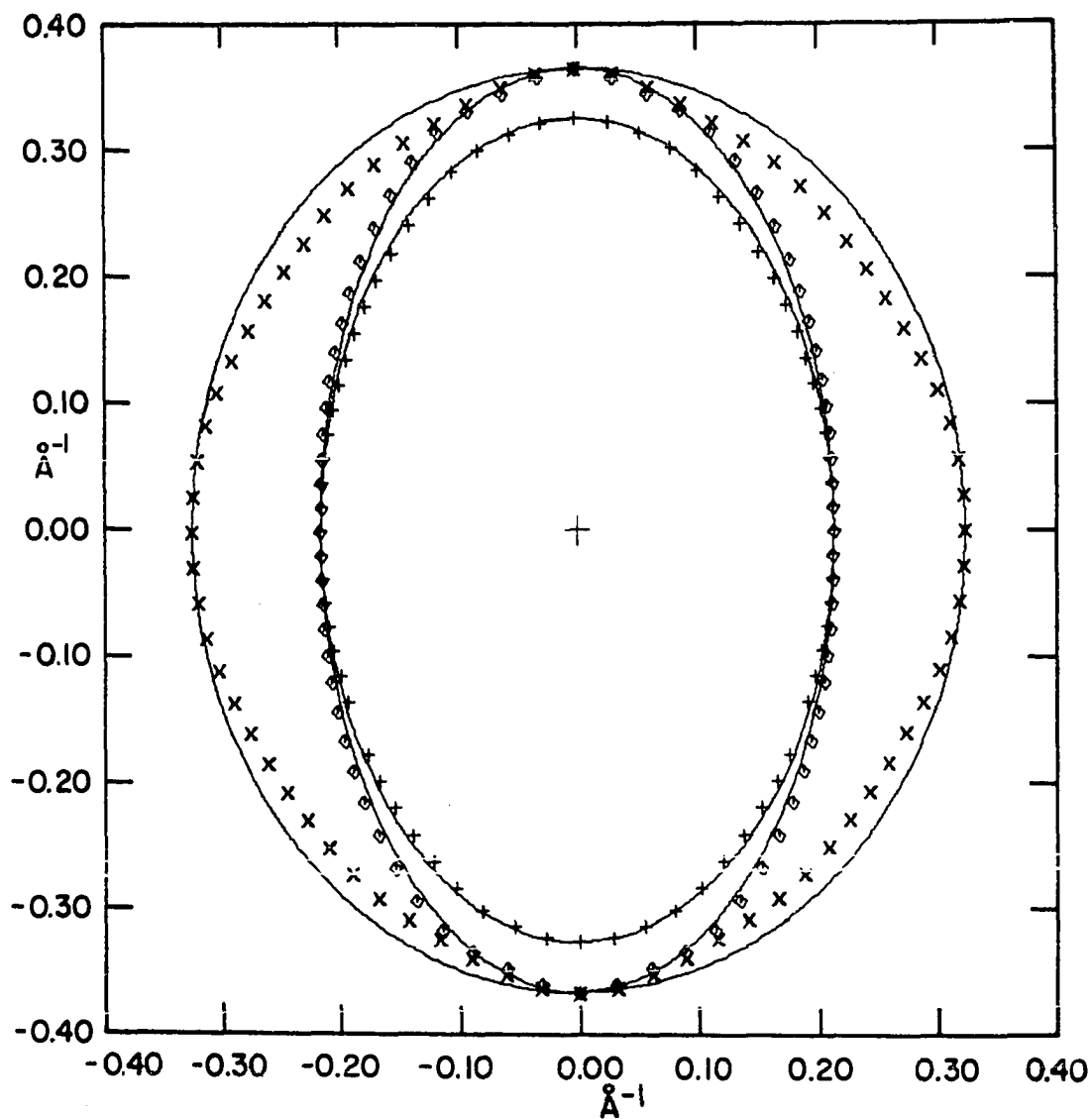


Figure 17. Cross sections of the hole ellipsoid Fermi surface at N for the planes containing the three semiaxes. These directions are NI , NH , and NP

inversion scheme for the $\ell = 16$ fit involving 45 coefficients. These 45 coefficients along with the two spherical mapping parameters, ALPHA and GAMMA, are listed in Appendix A. The $\ell = 16$ fit was the highest-order fit that could be obtained and was chosen because the standard deviation of the input data from that predicted by the inversion scheme was lower than that of any other fit obtained for a lower ℓ value.

Values for the three semi-axes of the ellipsoids at N as found with the inversion scheme are listed in Table 3. Since it is impossible to decide upon the assignment of semi-axes to the NH and $N\Gamma$ directions from the dHvA data, we assume that the shortest semi-axis is along NH, which is consistent with the FS calculations for Mo by Loucks (4). The volume calculated using the inversion coefficients in a suitable program from Reference 34 was found to be 0.614 \AA^{-3} for 6 hole ellipsoids. This compares with 0.61 \AA^{-3} obtained by Boiko et al. (18) and 0.71 \AA^{-3} obtained by Sparlin and Marcus (16).

Table 3. Values in \AA^{-1} for the semi-axes of the hole ellipsoids at N

| Direction | Inversion scheme radii | Estimated from three areas | Other investigations | | |
|-----------|------------------------------|-------------------------------------|-----------------------------|---------------------------------|-----------------------------|
| | | | Boiko <u>et al.</u> (18) | Cleveland & Stanford (20) | Sparlin & Marcus (16) |
| NP | 0.365 | 0.358 | 0.38 | 0.35 | $0.39 \pm .01$ |
| $N\Gamma$ | 0.325 | 0.310 | 0.29 | 0.32 | $0.30 \pm .01$ |
| NH | 0.216 | 0.223 | 0.22 | 0.20 | $0.23 \pm .01$ |

The column labelled "estimated from three areas" in Table 3 contains semi-axes as calculated from the frequencies (and corresponding areas) which exhibit the condition that $dF/d\theta = 0$ at $\langle 110 \rangle$ or $\langle 100 \rangle$ directions. These three frequencies correspond to the extremal cross sections π_{ab} , π_{bc} , π_{ac} for a perfect ellipsoid at N. The semi-axes a, b, and c calculated in this manner can only be as good as the approximation that the actual surface is a perfect ellipse. As we can see from Figure 17, the largest cross section has some places which are flatter than the perfect ellipse with same semi-axes, so we would expect semi-axes estimated in this manner to differ from those determined by the inversion scheme. The inversion scheme radii for this surface should be accurate to about 1%.

D. Electron Lenses at Δ

Data related to the electron lenses at symmetry point Δ along ΓH in the Brillouin zone cannot be inverted by Mueller's inversion scheme, since this point does not possess inversion symmetry. The physical significance of this fact is that dHVA experiments cannot provide enough information about this surface to specify a unique surface consistent with the required C_{4v} symmetry. If an additional condition is imposed that the surface has a mirror plane whose normal is parallel to ΓH , the point has now been given the D_{4h} symmetry, which satisfies inversion theorem requirements, and enables use of the Mueller inversion scheme or use of a geometrical model with the additional symmetry. However, we may only obtain a possible substitute for the actual surface when we impose this additional requirement to obtain a unique fit to the data. It is possible, but rather unlikely, that the real FS exhibits this mirror symmetry.

The condition of four-fold symmetry in a plane whose normal is parallel to ΓH for the point Δ and the almost zero slope of the lowest frequency branch of the experimental data for the (001) plane (see Figure 7) can be combined to suggest that the lenses are very nearly circular with radius r in a cross section whose normal is parallel to ΓH . The dHvA data about the lens' cross section having a dimension d along ΓH as one axis and the lens' radius r as the other axis is not sufficient to specify the distance α between the center of the plane of the circular cross section and the center of the ΓH dimension. It can be shown by simple geometrical arguments that two different cross sections with identical values for r and cross-sectional areas, but differing in values of α and d should be distinguishable because of differing caliper dimensions for the same angle of observation.

This effect is probably only noticeable as a 4-5% difference in caliper dimensions, unless the lenses are highly asymmetrical. In other words, the RFSE might be able to distinguish between symmetrical and asymmetrical models which could both agree with the dHvA results. However, unless the RFSE data is very accurate, probably better than 1%, the amount of asymmetry along the ΓH direction could not be determined quantitatively. Such accuracy can not be expected from RFSE orbits yielding small caliper dimensions, since these resonances are detected at low magnetic field strengths, where $\omega_c \tau$ is low and the RFSE line shapes are broad. Another result of the geometrical arguments is that while the distance α can vary quite widely, the ΓH dimension for the lens remains reasonably constant, as does the volume of the FS.

For the above reasons, we will use the five frequency branches for the (001) and (1 $\bar{1}$ 0) planes shown in Figure 7 as input to the inversion scheme for the D_{4h} symmetry group. As with the ellipsoids discussed in Section III-C, we pick one lens and place it at the origin of our inversion k-space and use all of the data as if it had been taken from five different planes of observation. The program was checked with perfect ellipsoids of revolution and the spherical mapping parameter was also determined from this ideal data.

The two principal cross sections for the lens pieces situated at symmetry points Δ are shown in Figure 18, subject to the additional symmetry requirement necessary for inversion of the data. The results obtained from the inversion scheme for the $\ell = 16$ fit involving 25 coefficients, which are listed along with the spherical mapping parameter GAMMA in Appendix A, came from the highest order fit that could be obtained.

Values for the radii as found with this fit are listed in Table 4.

Table 4. Values in \AA^{-1} for the radii of the electron lenses at Δ

| Direction | Inversion scheme radii | Other investigations | |
|---------------------------------------|------------------------------|----------------------|------------------------|
| | | Boiko et al. (18) | Brandt & Rayne (15) |
| For \sim circular cross section: | | | |
| parallel to $\langle 100 \rangle$ | .162 | .155 | .16 |
| parallel to $\langle 110 \rangle$ | .155 | | |
| Total TH dimension | .222 | .22 | .22 |

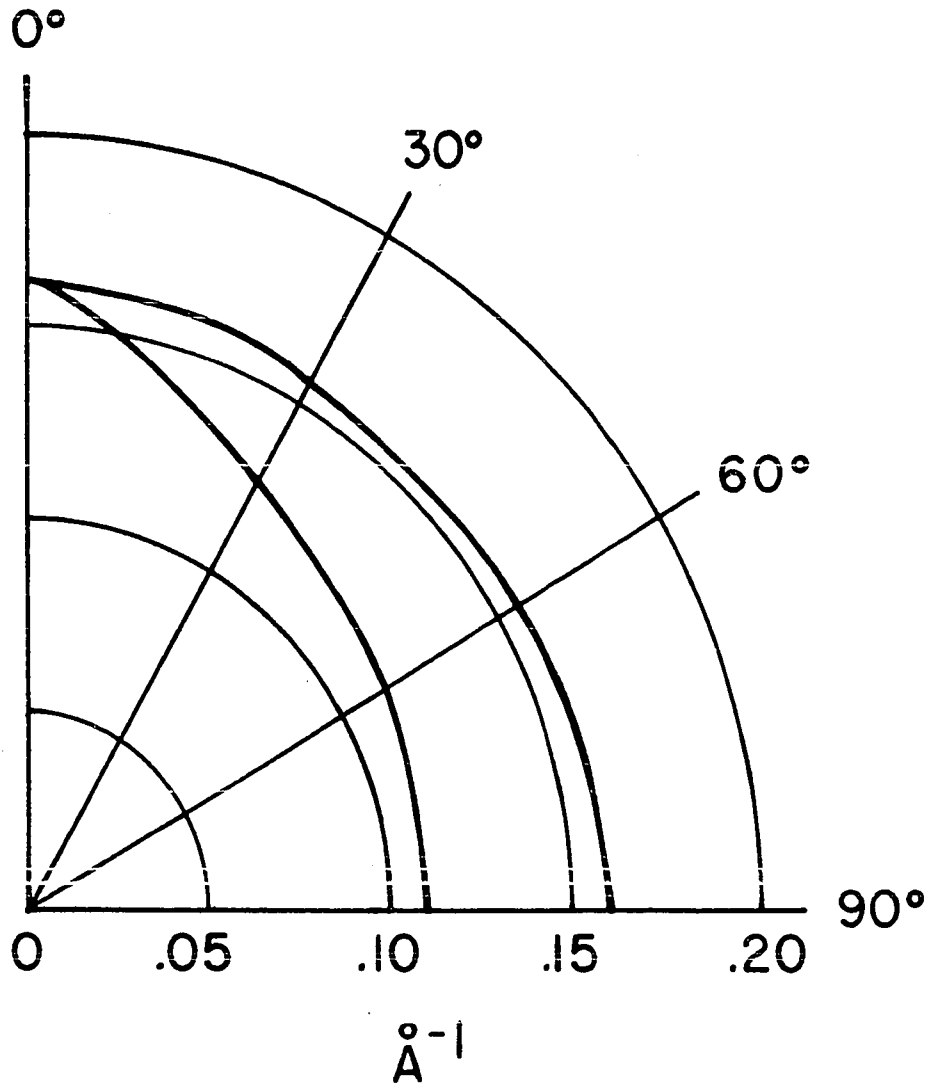


Figure 18. Cross sections of the electron lens Fermi surface at Δ assuming a higher D_{4h} symmetry. For the nearly circular cross section, the plane has normal direction ΓH , and 0° and 90° are parallel to $\langle 100 \rangle$ directions

The volume calculated using the inversion coefficients in a suitable program from Reference 34 was found to be 0.057 \AA^{-3} for 6 lenses. This compares with a volume of 0.060 \AA^{-3} obtained by Leaver and Myers (17).

E. Electron Jack at Γ

An accurate model for the electron jack surface is the most difficult to extract from the dHvA data of extremal cross-sectional areas. The data can not be inverted by using the Mueller inversion scheme because some radius vectors to points on the surface are multi-valued. There is no way to accurately determine the position of the center of the ball connected to the body portion of the jack. In Section III-B we have seen that the RFSE data of Boiko et al. (18) agree quite closely with the radii obtained from inversion of the octahedron dHvA data. Their RFSE data will be used to supply some very valuable jack radii, but only at those angles where there is good reason to believe that the caliper dimensions obtained in the RFSE experiment are actually radii. The claim made by Boiko et al. (18) that "upon rotation of the magnetic field relative to the crystallographic directions in the (110) plane, the lines of the experimental points from the extremal orbits of the electron 'jack' will give directly the section of this surface with the (110) plane" is not true for the entire jack cross section in this plane. In particular the line i of the data of Boiko et al. (18), shown as plotted points in Figure 19, can be coincident with the actual radius vector only at the [001] direction. At any other angular position for this line, the RFSE measures caliper dimensions, and must be interpreted as such. It is easy to fall into the trap of equating caliper dimensions with radii for

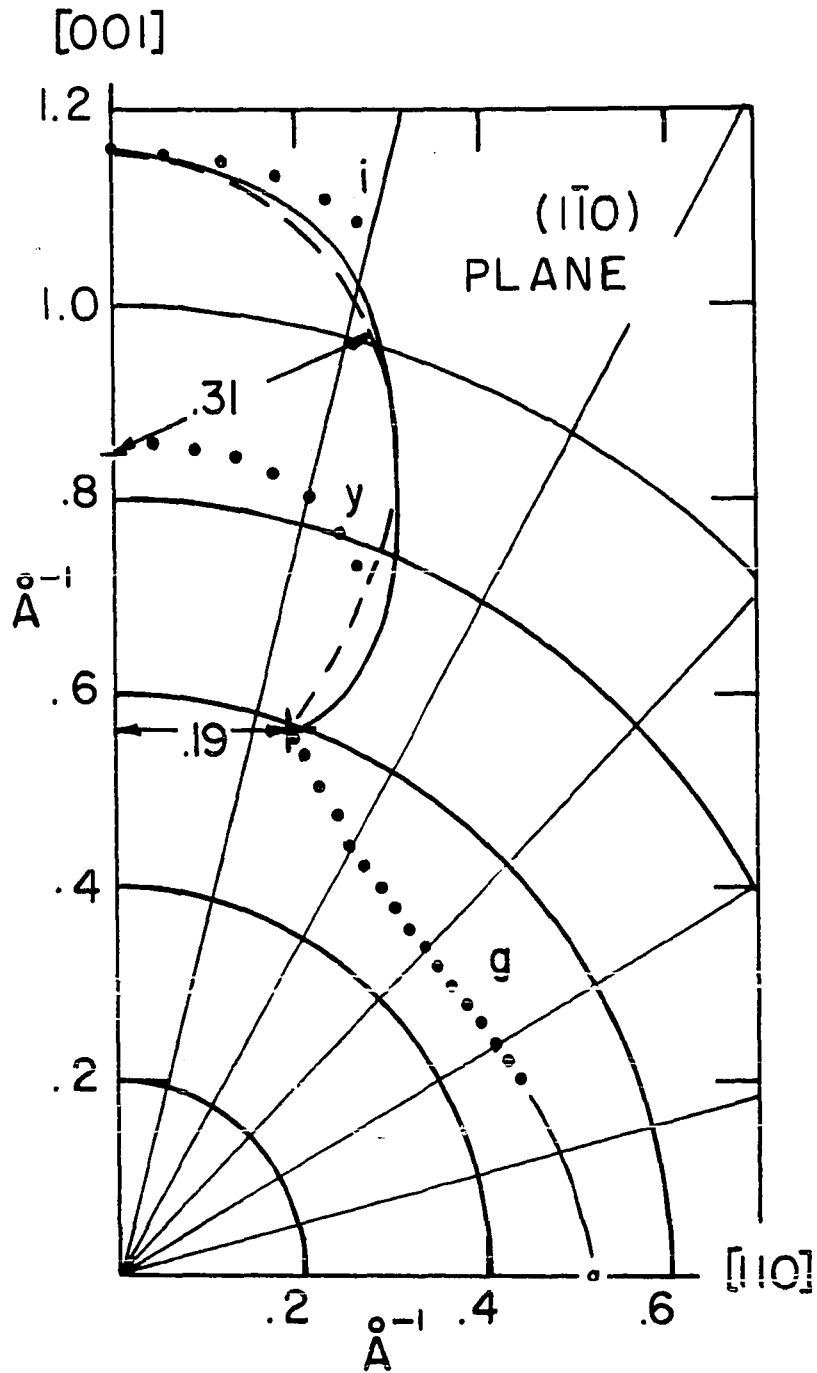


Figure 19. $(1\bar{1}0)$ plane cross section for the electron jack Fermi surface at Γ . The solid and open circle data points and their labels are RFSE results of Boiko *et al.* (18). The + data point at 19° from $[001]$ and the solid line construction are described in the text

surfaces whose curvature and symmetry does not allow such a statement to be true. In general extreme care should be exercised when choosing the RFSE calipers which also have the property of being FS radii.

In the following paragraphs the rationale for the particular $(1\bar{1}0)$ and (001) cross sections for the electron jack as shown in Figures 19 and 20 will be presented. First some facts obtained directly from our data should be noted. In Section III-D we saw that the electron lenses are very nearly circular. Since the lenses lie in the neck portion of the jack and the lens and neck cross sections are comparable in size, it is reasonable to assume that the necks are also very nearly circular. From the value of 11.7 MG at the $\langle 100 \rangle$ directions for the neck orbit labelled σ in Figure 6, we obtain a radius of 0.19 \AA^{-1} for a circular neck cross section. Looking at the $(1\bar{1}0)$ plane for the ball orbit labelled π in Figure 6, we see values of 31.63 MG at $[001]$, 33.4 MG at $[110]$, and 35.4 MG at $[111]$. We also note that the neck cross section is 32.4 MG at $[110]$. Assuming a rather spherical shape for the ball, and a circular cross section for the $[001]$ direction, we calculate a radius of $.310 \text{ \AA}^{-1}$ for the ball. The angles at which the body orbit labelled ξ in Figure 6 disappear in the $(1\bar{1}0)$ plane can help to define the neck dimensions. At a magnetic field direction of $19 \pm \frac{1}{2}^\circ$ from $[110]$ the neck interferes with the completion of the body orbit, and the orbit is no longer possible. This corresponds to the loss of RFSE signals at $21 \pm 2^\circ$ from a $\langle 100 \rangle$ direction reported by Cleveland and Stanford (20), and at $21 \pm 3^\circ$ from a $\langle 100 \rangle$ direction reported by Boiko et al. (18). Note that RFSE data are plotted as caliper dimensions for given directions and are obtained with

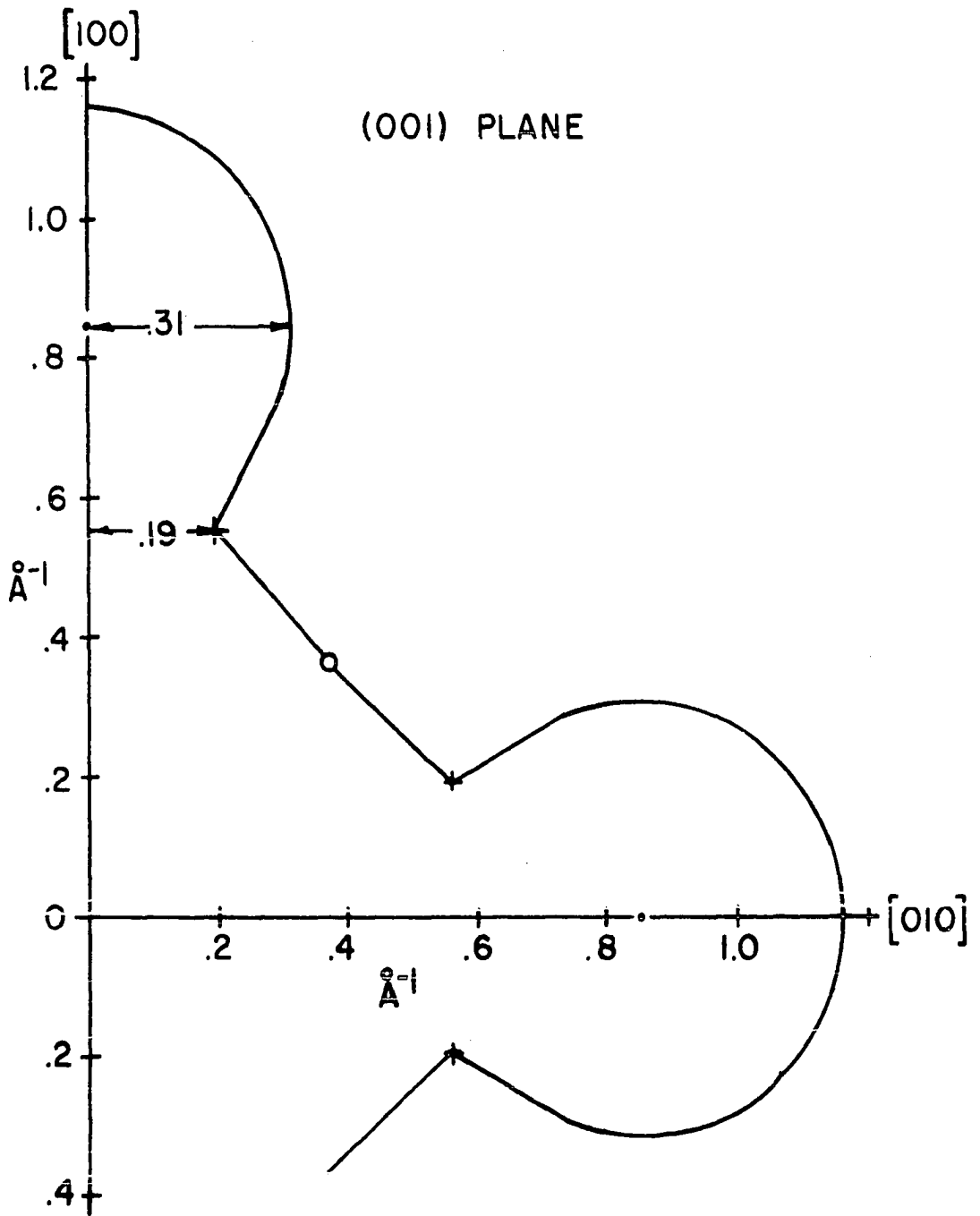


Figure 20. (001) plane cross section for the electron jack Fermi surface at Γ . The open circle is a data point from Reference 18. The + data points and the solid line construction are described in the text

fields applied perpendicular to the given direction and to the plane of the sample. At a magnetic field direction of $27 \pm 2^\circ$ from $[100]$, the orbit around the body changes into an orbit which encompasses all four balls. A loss of RFSE signals at $29 \pm 2^\circ$ from a $\langle 110 \rangle$ direction was reported by Cleveland and Stanford, and at $24 \pm 3^\circ$ from a $\langle 110 \rangle$ direction by Boiko et al.

In Figure 19 for the $(1\bar{1}0)$ plane we assume the RFSE calipers of Boiko et al. at $[001]$ and $[110]$ are radii, as well as those arising from the body of the jack, labelled g. The radii are 1.16 \AA^{-1} at $[001]$ and $.52 \text{ \AA}^{-1}$ at $[110]$. We extrapolate an additional data point + for the body at 19° from $[001]$ which is indicated by our data. This point also defines the neck radius at 0.19 \AA^{-1} for a circular neck cross section. A completely spherical ball is shown dashed in along with a straight line to connect the neck to the ball, and a gently curved solid line is drawn connecting the body data with the data point at $[110]$. We shall return to this figure shortly.

We continue to assume a spherical ball with circular neck and ball cross sections at $\langle 100 \rangle$ directions for the ball and neck drawn in Figure 20 for the (001) plane. The RFSE data point at $[110]$ is included and a slightly curved line is drawn to connect the necks with this point. The body is slightly concave at $[110]$ in this plane, which agrees with the observation of two neck frequencies near 32.4 MG at small angles away from $[110]$ in the (001) plane, as shown in Figures 6 and 8. The area computed for the model (001) plane is 2.19 \AA^{-2} which compares very favorably with 2.21 \AA^{-2} for the dHvA experimental area.

An area integration of the $(1\bar{1}0)$ plane as we have constructed it thus far yields 1.45 \AA^{-2} to be compared with the dHvA experimental value of 1.51 \AA^{-2} . We can improve the value for this area and also correct the model so that both a minimum (neck) and maximum (ball) cross section is observable at the $[110]$ direction where the experimental data shows values of 32.4 MG and 33.4 MG by drawing the additional (solid) lines for the ball. The new area for the $(1\bar{1}0)$ plane is 1.49 \AA^{-2} . An integration to check the neck and ball cross sections at $[110]$ quantitatively is impossible to do accurately without a complete analytical description of the surface to be analyzed.

A possible cross section for the $(1\bar{1}1)$ plane is shown in Figure 21. The radii at the $\langle 110 \rangle$ and $\langle 112 \rangle$ directions are the same as indicated for similar directions in Figures 19 and 20. The model area of 0.77 \AA^{-2} can be compared with the experimental value of 0.81 \AA^{-2} for the $[111]$ direction.

It is not possible to carry this simple geometrical model further or make additional comparisons with the experimental data since the integrations of cross-sectional areas which do not contain the center of symmetry for the figure and/or whose normal is not a direction of high symmetry are possible only with an analytical model suitable for a computer. The assumption about fairly spherical balls is also likely to break down somewhat as one adjusts the model for a more accurate fit. As mentioned in Section III-A, the analytical approach was tried with no better results than those presented above.

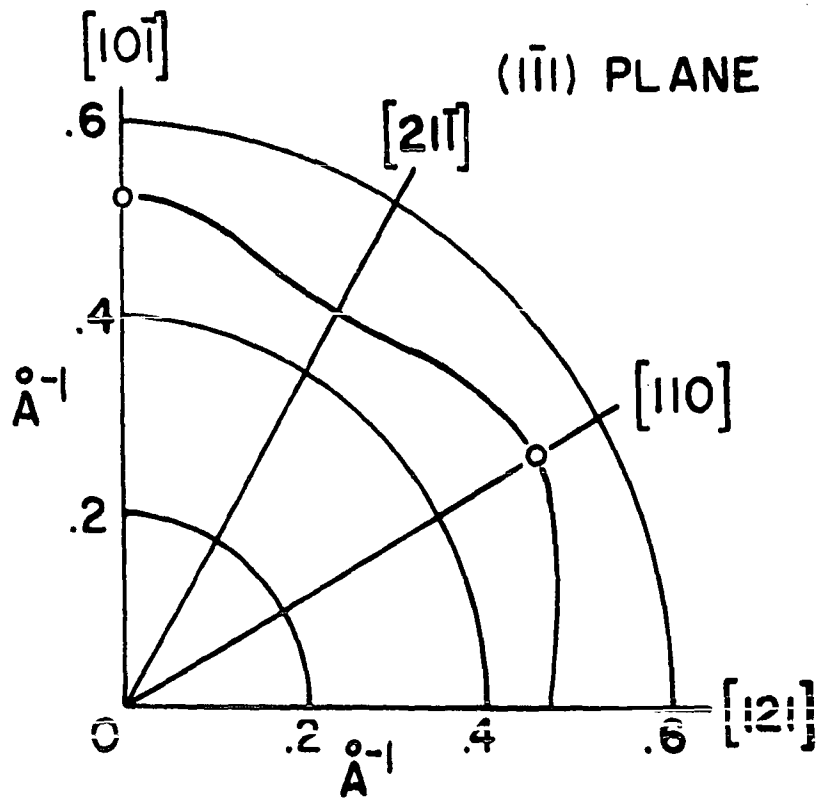


Figure 21. A possible $(1\bar{1}\bar{1})$ plane cross section for the electron jack Fermi surface at Γ . The open circles are data points from Reference 18

IV. CONCLUSIONS

We have presented dHvA data for molybdenum which are accurate to within 1% in frequency or 1° in angle, whichever is the larger error, for the magnetic field rotated in the $(1\bar{1}0)$, (001) and $(1\bar{1}1)$ crystallographic planes. Two samples, cut from the same single crystal rod as the sample used in the RFSE work of Cleveland and Stanford (20), were used to obtain the dHvA data. These data were shown to be consistent in frequency and angular determination at the appropriate major symmetry directions for the above-mentioned planes. The data were inverted through the use of the Mueller inversion theorem (35), when possible, to obtain FS radii, and a simple geometrical model was developed for the electron jack. The FS radii for the $(1\bar{1}0)$ plane cross section for the hole octahedron obtained from the dHvA data were shown to be consistent with the RFSE caliper dimensions obtained by Boiko et al. (18) for the same plane. The geometrical model for the electron jack cross section is consistent with the RFSE caliper dimension at $[001]$. Since the inversion radii data for the octahedron are most subject to error at $[001]$, we believe that the determination by Boiko et al. of the RFSE caliper dimensions, which are radii at $[001]$, to be correct. These RFSE dimensions for the $\langle 100 \rangle$ directions are 1.16 \AA^{-1} for the jack and 0.79 \AA^{-1} for the octahedron with the ΓH dimension as 1.999 \AA^{-1} , so the gap between the jack and octahedral FS pieces is estimated as 2.5% of the ΓH dimension, thus resolving the discrepancies existing in current RFSE data.

The model presented for the electron jack could probably be improved

upon by further calculations on a complicated analytical model. However, an accurate FS model based on band structure calculations for molybdenum would be more desirable. Such a calculation has yet to be completed by band theorists.

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To my wife, Bea, go special thanks for her patience, understanding, and constant encouragement during these past years of study and research.

The secretarial skills of Mrs. Carol Kline are appreciated in the typing of the manuscript.

VII. APPENDIX A

COEFFICIENTS FROM INVERSION SCHEME FITS

| HOLE OCTAHEDRON | | | ELECTRON LENS | | |
|-----------------|----|---------------------------|---------------|----|---------------------------|
| # | J | RADIUS**2 COEFFICIENTS | # | J | RADIUS**2 COEFFICIENTS |
| 1 | 0 | 0.131200D 03 | 1 | 0 | 0.879556D 01 |
| 2 | 4 | 0.300859D 02 | 2 | 2 | 0.448323D 00 |
| 3 | 6 | -0.144265D 01 | 3 | 4 | 0.634310D 00 |
| 4 | 8 | 0.940868D 01 | 4 | 4 | 0.855809D-01 |
| 5 | 10 | -0.387562D 00 | 5 | 6 | -0.655377D-01 |
| 6 | 12 | 0.242699D 01 | 6 | 6 | -0.310930D-01 |
| 7 | 12 | 0.128054D 00 | 7 | 8 | 0.944097D-01 |
| 8 | 14 | -0.643953D-01 | 8 | 8 | 0.147971D-01 |
| 9 | 16 | 0.161086D 01 | 9 | 8 | 0.253498D-02 |
| 10 | 16 | 0.267312D 00 | 10 | 10 | -0.176735D-01 |
| 11 | 18 | 0.127289D 00 | 11 | 10 | -0.183491D-01 |
| 12 | 18 | 0.644677D 00 | 12 | 10 | 0.738371D-02 |
| 13 | 20 | 0.447419D 00 | 13 | 12 | 0.261165D-01 |
| 14 | 20 | 0.393513D 00 | 14 | 12 | 0.320903D-01 |
| 15 | 22 | -0.155024D 00 | 15 | 12 | 0.824302D-02 |
| 16 | 22 | 0.275068D 00 | 16 | 12 | 0.560084D-02 |
| 17 | 24 | 0.797590D 00 | 17 | 14 | -0.839417D-02 |
| 18 | 24 | -0.198070D 00 | 18 | 14 | 0.350407D-02 |
| 19 | 24 | 0.187362D 00 | 19 | 14 | -0.386925D-02 |
| | | | 20 | 14 | -0.104948D-01 |
| | | | 21 | 16 | 0.893395D-02 |
| | | | 22 | 16 | 0.129436D-01 |
| | | | 23 | 16 | -0.905301D-02 |
| | | | 24 | 16 | -0.953375D-02 |
| | | | 25 | 16 | 0.114243D-01 |

GAMMA = 1.6078

HOLE ELLIPSOID

| # | J | RADIUS**2 COEFFICIENTS | # | J | RADIUS**2 COEFFICIENTS |
|----|----|---------------------------|----|----|---------------------------|
| 1 | 0 | 0.3528140 02 | 23 | 12 | -0.1859230-01 |
| 2 | 2 | -0.8318530-02 | 24 | 12 | -0.6649740-01 |
| 3 | 2 | -0.5975690 00 | 25 | 12 | 0.1466700-01 |
| 4 | 4 | 0.2107760 00 | 26 | 12 | -0.5174170-01 |
| 5 | 4 | 0.1020840 01 | 27 | 12 | -0.6018350-02 |
| 6 | 4 | 0.2669030 00 | 28 | 12 | 0.1268780-01 |
| 7 | 6 | 0.1247220-02 | 29 | 14 | 0.4813400-01 |
| 8 | 6 | -0.6776950-01 | 30 | 14 | -0.6223770-02 |
| 9 | 6 | -0.7651980-01 | 31 | 14 | -0.4132820-01 |
| 10 | 6 | 0.5822000-02 | 32 | 14 | 0.3114650-01 |
| 11 | 8 | 0.4492170-01 | 33 | 14 | 0.2541590-01 |
| 12 | 8 | 0.2564740-01 | 34 | 14 | -0.2756810-01 |
| 13 | 8 | 0.5237700-01 | 35 | 14 | -0.2596160-01 |
| 14 | 8 | 0.2310100-02 | 36 | 14 | -0.1786960-01 |
| 15 | 8 | 0.1784030-01 | 37 | 16 | 0.1748390-01 |
| 16 | 10 | 0.1532800-01 | 38 | 16 | -0.2231940-01 |
| 17 | 10 | 0.9287760-02 | 39 | 16 | -0.4286940-01 |
| 18 | 10 | -0.1132440-01 | 40 | 16 | 0.1740480-01 |
| 19 | 10 | 0.1752080-01 | 41 | 16 | -0.7478760-02 |
| 20 | 10 | 0.3763650-04 | 42 | 16 | -0.2238790-01 |
| 21 | 10 | 0.3534310-01 | 43 | 16 | -0.3007020-01 |
| 22 | 12 | 0.6050980-01 | 44 | 16 | 0.3235280-01 |
| | | | 45 | 16 | 0.9382680-02 |

ALPHA = 1.3904
GAMMA = 0.8638

VIII. APPENDIX B

FREQUENCY ANALYSIS PROGRAM LISTING AND DISCUSSION

This appendix describes the computer programs used to transfer and analyze the dHvA data. Job A334TTT, "TAPE7 TO TAPE9," transfers the data from a 7-track magnetic tape produced from the punched paper tape output created at the time the experimental equipment was being operated to a 9-track tape for data analysis and storage. Job B334FFT is the program responsible for the data analysis. Both of these listings appear after the discussion part of this appendix.

In "TAPE7 TO TAPE9" the assembler language routine CHGTC takes the intermixed dHvA signal and magnetic field value arrays which are input as 4096 characters in the array Y, and maps these characters in a 1:1 fashion to a new character set which is input for the assembler language routine TRANS. TRANS decodes the Y character array and returns an array Y containing 2008 data points and an array H containing 20 values for the magnetic field. The field values are then converted from integer values ranging from 0 to 4095 to actual readings in megagauss through the use of the constants set at the time the experiment was conducted. An identification card, the Y and H arrays, and the experimental constant data are recorded on 9-track tape as pairs of records for each data set to be read later by the data analysis program. Appropriate labels and information about the magnetic field values are printed to enable one to determine whether the data set has been converted and stored correctly. A plot of typical data for an angle of 28.9° from [010] in the (001) plane is shown in Figure 22. The Y array of integer values (up to ± 128) of the dHvA

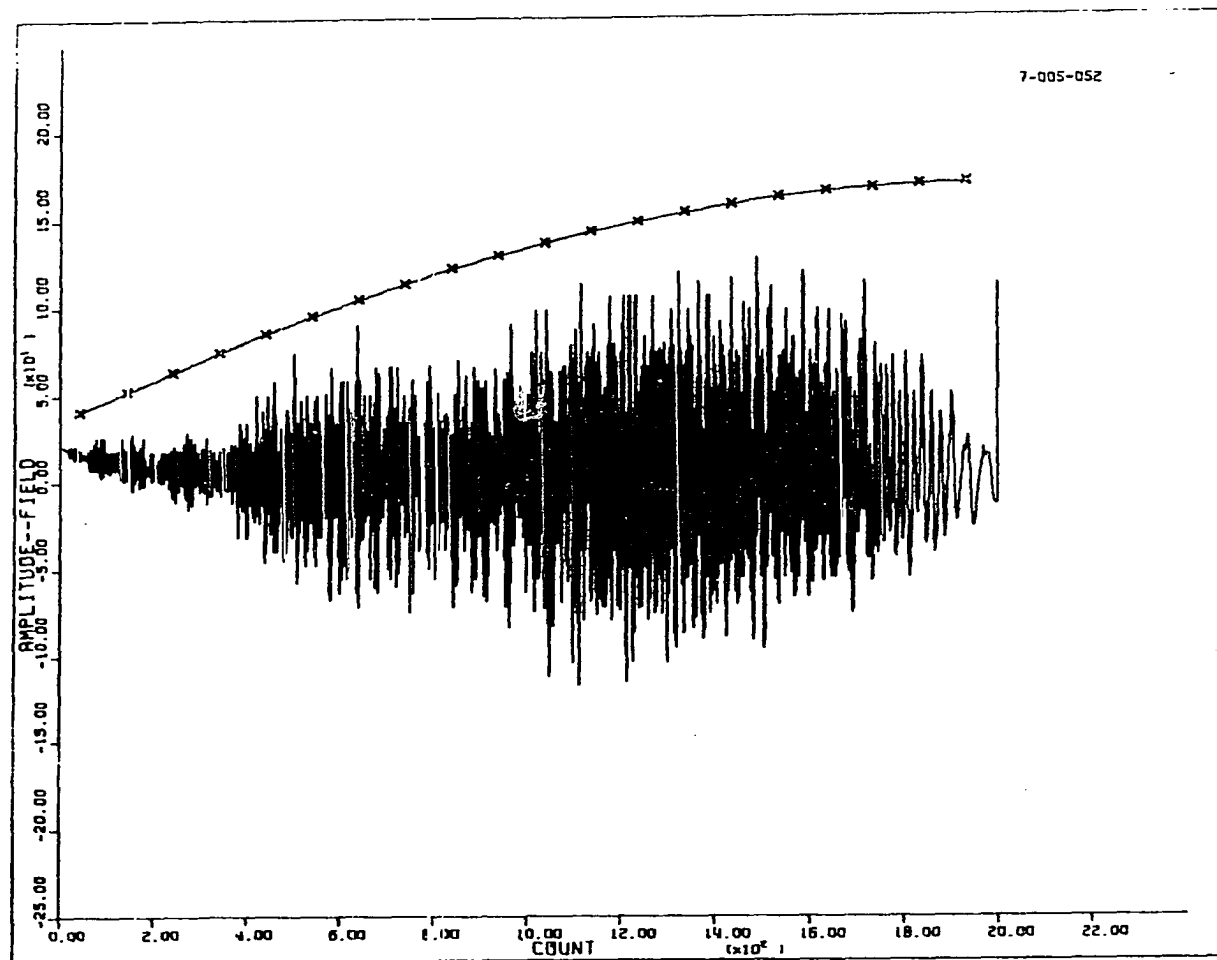


Figure 22. A plot of typical dHvA data. This data was taken at 28.9° from [010] in the (001) plane

signal was plotted versus integer count of time in units of $4 \mu\text{sec}$. The H array of field values is in units of megagauss for the same integer time count.

The analysis program B334FFT used the fast-Fourier subroutine RHARM from the IBM Scientific Subroutine Package. This analysis is 10 - 60 times faster than the filter-periodogram technique described by Panousis (29) with comparable frequency accuracy and resolution. Figure 23 shows a filter-periodogram frequency analysis of dHvA data taken at 28.9° from [010] in the (001) plane and shown in Figure 22. Figure 24 shows the frequency analysis of the same data set by the fast-Fourier technique. The peaks labelled β_1 , β_2 , β_3 are plotted as points labelled β for frequencies at 28.9° from [010] in the (001) plane as shown in Figure 6. Peaks labelled with subscripts in Figure 24 (such as $(\beta_1)_2$) are second harmonics, and peaks with two frequencies (such as $\gamma \pm \pi$) are sums and differences of the appropriate fundamental terms. Only frequencies identified as fundamental were plotted in Figure 6. In both Figures 23 and 24 the graphs are composites of several small-range frequency scans, since each program does its best on frequency scans of one octave or less. For purposes of comparison the frequency ranges and selected data ranges used were identical for both analyses. The computer time used to obtain the filter-periodogram frequency results for the data set was around 35 minutes of CPU time, while the same frequency results were obtained by B334FFT with about 45 seconds of CPU time. From the 27 peaks identified in Figure 24 we get some feeling of the power of either analysis program to extract many frequencies with high accuracy and resolution from a single data set.

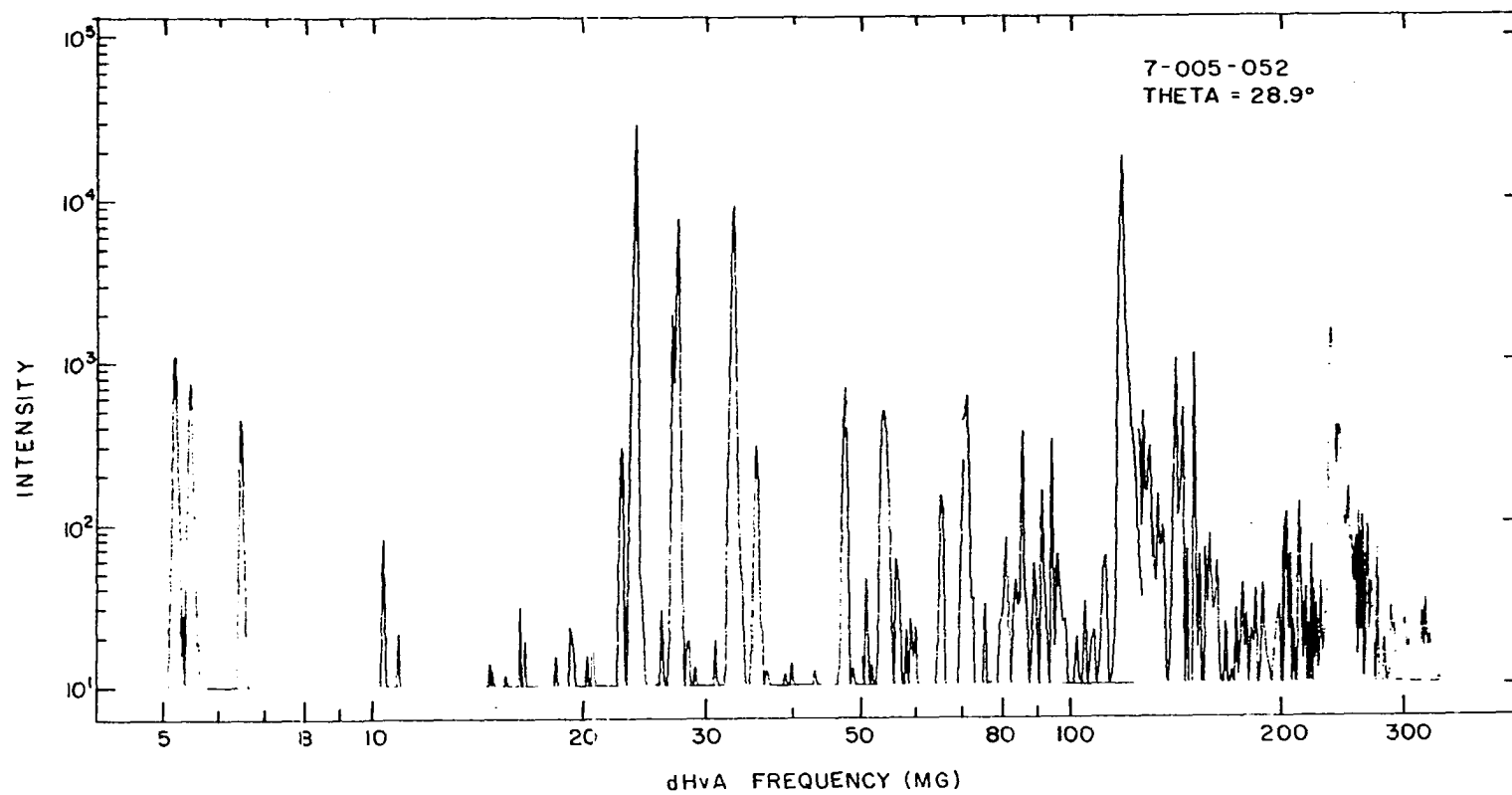


Figure 23. A filter-periodogram frequency analysis of the dHvA data shown in Figure 22

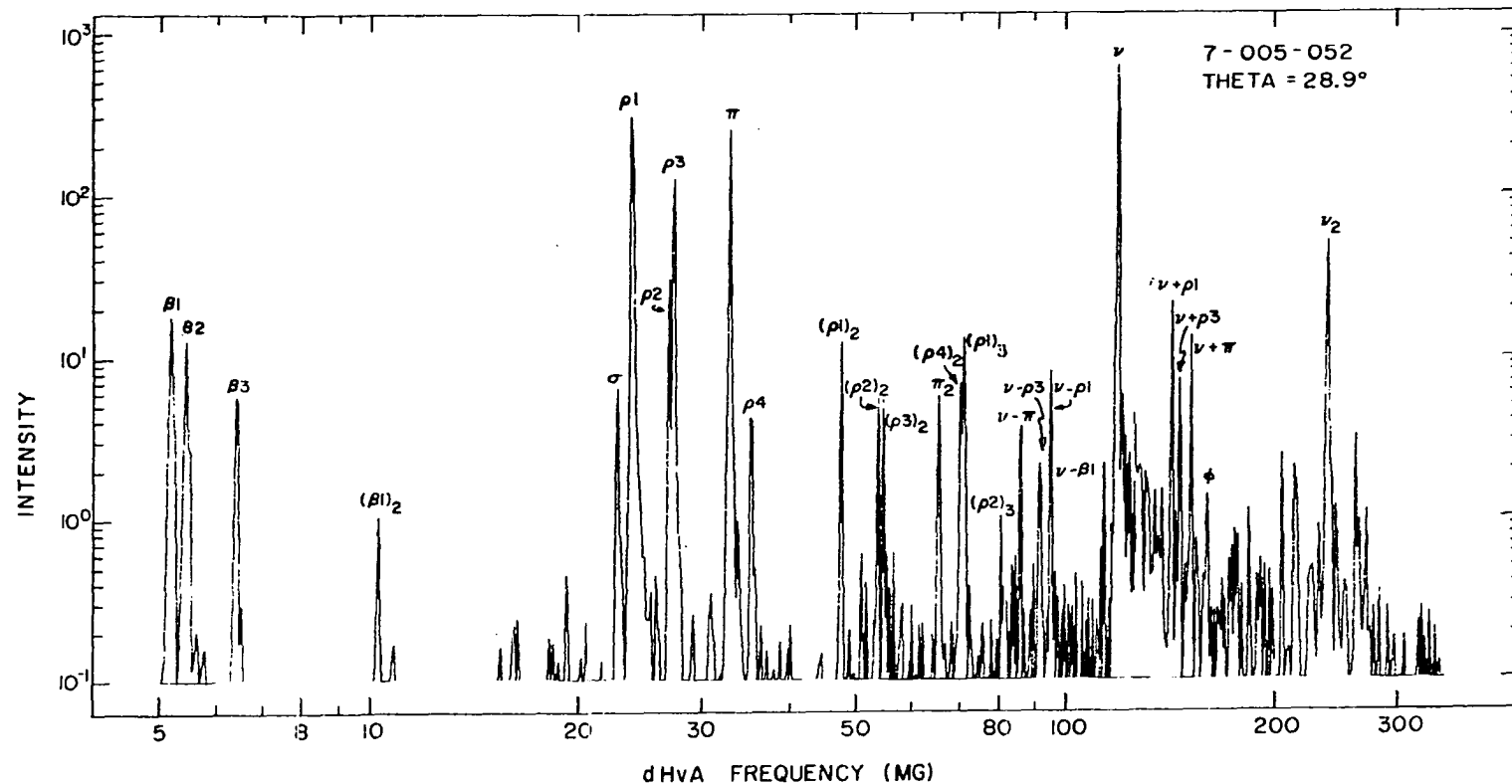


Figure 24. A fast-Fourier frequency analysis of the dHvA data shown in Figure 22. The fundamental frequencies from this analysis are plotted and labelled in Figure 6 at 28.9° from $[010]$ in the (001) plane

The intensity scales on Figures 23 and 24 should not be compared for absolute intensity.

B334FFT was written with several options to handle the most frequently desired modes of data input and analysis. It assumes that the data is identified by the same cards used with A334TTT to create the 9-track data storage tape. Some named COMMON storage areas are used to pass different variables at different stages in the calculation and/or used for work space to keep storage space to a minimum. The program will run in either a (160K,32K) region or (96K,96K) region of fast-core and bulk-core storage depending upon how the HIRARCHY card is set up in the LKED step of the executed procedure FORTH.

All reading of tapes and cards is done in subroutine READIN. BPASS is equal to 0.0 for the first call of this subroutine for initialization of graph labels and program options. A set of frequency cards is read in and stored, a blank card is detected, followed by the first data identification (ID) card which is read. The appropriate data set is obtained from a sequential tape and control is returned to MAIN. Each further call to READIN reads an ID card and provides an additional data set for analysis. If a blank card is encountered, a new set of frequencies, one blank card, and one ID card are again read in before control is returned to MAIN.

Subroutine BFIELD corrects for the time delay effects described by Equation 4. For all of this investigation TZERO=40 μ sec was found appropriate to obtain approximately equal frequency results for both "rising field" and "falling field" data. Since molybdenum is not ferromagnetic,

HZERO=0.0 for this work.

Subroutine FITH identifies the type of field contained in the data as "rising" and/or "falling", checks that the number of data points associated with each identification is acceptable, and returns the code of the type of field detected, the appropriate calculated array of up to 2008 field values, and the limits on the array.

Subroutine RANGE takes into account the fact that the amplifiers in our apparatus are capable of passing time frequencies in only a certain given range ($\sim 5 - 100$ kHz), and that our dHvA digital equipment is capable of taking data at equal time intervals of $4 \mu\text{sec}$ minimum (i.e., only two points per cycle are obtained for a 125 kHz signal). Since dHvA oscillations are periodic in $1/H$ rather than in real time, for a given pair of data points there is a minimum and a maximum value for possible dHvA frequencies contained in the data, as given by the relation

$$\text{time freq.} \doteq \left| \frac{\Delta\left(\frac{1}{H}\right)}{\Delta T} \right| \text{ dHvA freq.} \quad (7)$$

with ΔT and ΔH computed from this pair of points and the time frequency being one of the amplifier frequency limits. The upper amplifier limit assures us of taking data at a rate of more than two points per cycle. We have found that when data is analyzed for a given dHvA frequency using only those data points which lie within the time frequency limitations given above, we obtain the best analysis.

This subroutine works best if the desired frequency ranges are input in increasing order, and computes starting and stopping counts for analysis of the data for each of the frequency ranges it is given. The

routine is bypassed by the MAIN program if the starting and stopping count values are overridden from data input cards.

Subroutine GENY converts a portion of the dHvA data array YI, taken in equal increments of time, to an expanded array Y of 16,384 data points computed at equal increments of $1/H$, since the dHvA data is periodic in $1/H$ rather than in time, by using the inverted field values in the HI array. The number of data points used in the conversion depends upon NSTART(I) and NSTOP(I) as usually determined by the RANGE subroutine for the Ith frequency range. Straight-line interpolation is used between data in both the YI and HI arrays and is generally a good approximation. Trying to increase the resolution in the data by setting RES \neq 0 on input cards and filling part of the large Y array with zeroes was not found to be a satisfactory means of increasing the resolution. It should be noted that this mapping from real time space to $1/H$ space causes time-periodic noise signals to be transformed into random noise in $1/H$ space.

Subroutine OUTPUT provides printed output of the frequency analysis and optional tape records and/or graphic display of the results. Tape output was used to input frequency results for computer averaging by another program. If the frequency range desired contains more than 300 possible output frequencies, the output is broken down into scans which do not overflow the output data arrays. Subroutine PEAKS is called by OUTPUT to search for peaks in the frequency spectrum and to fit a second order curve to determine the true peak frequency and its intensity. These frequencies are listed by OUTPUT if their intensities are greater than .005 or any value set by the override variable ALIM.

The MAIN routine sets up the actual frequency ranges to be analyzed and controls the appropriate choices of subroutines for each particular analysis, depending on information generated from the previous analysis. The original data set is stored and restored, if needed, by this routine.

The IBM subroutine RHARM analyzes $N = 2 \cdot 2^M$ data points and returns $N+2$ Fourier coefficients corresponding to the constant coefficient and the sine and cosine coefficients on the first 2^M harmonics in the equation

$$x_j = \frac{a_0}{2} + (b_0=0) + \sum_{k=1}^{N-1} (a_k \cos(\frac{\pi j k}{N}) + b_k \sin(\frac{\pi j k}{N})) + \frac{a_N}{2} (-1)^j + (b_N=0) . \quad (8)$$

Test data run with this subroutine has led us to conclude that:

- (1) RHARM is capable of determining a single sine-wave frequency to the nearest integer $k \pm 0.5$.
- (2) Resolution of frequencies less than 3 cycles apart is generally not possible.
- (3) The frequency determined is not extremely dependent on the shape of a long-wave envelope containing the sine wave.
- (4) The frequency determined by RHARM does not appear to depend on the value of M , as long as the frequency determined is not approaching the highest frequency capable of being determined with a given M .
- (5) Phase shifting part of a wave train with respect to another part with identical frequency does cause severe problems in frequency determination. Many harmonics are necessary to reproduce the wave-form created, and the fundamental frequency in the data may

be smeared or be lost all together.

The intensity for the k^{th} harmonic in Equation 8 is computed from the Y array of returned coefficients of RHARM by using

$$\text{intensity} = (a_k)^2 + (b_k)^2 = (Y(n))^2 + (Y(n+1))^2 \quad (9)$$

for the frequency (in megagauss)

$$\text{frequency} = \frac{k}{\left| \frac{1}{H_1} - \frac{1}{H_2} \right|} = \frac{n-1}{2} \cdot \frac{1}{\left| \frac{1}{H_1} - \frac{1}{H_2} \right|} \quad (10)$$

where $n = 2k + 1$ is the index on the Y array. Note that n is always odd. If this is not true, incorrect pairs of coefficients from two different harmonics would be used in computing the intensity with incorrect results.

A listing of the programs and some comments on the called library subroutines follows.

```

//A334TTT  JOB  'A0367,SIZE=96K,TIME=4',JIM,MSGLEVEL=1
//STEP1 EXEC  FORTGCG,PARM.FORT='DECK',REGION.GJ=80K,TIME.GO=2
//FORT.SYSIN DD *
C      THE MAX. NUMBER OF DATA SETS ON TAPE7 IS 100.
C      NDSET = THE NUMBER OF DATA SETS TO BE READ.
C      NFSET = THE NUMBER OF DATA SETS PER FILE = 50, IF NOT SPECIFIED.
C      RATE,HSCALE,HZERO,THETA,COILK ARE SAME FOR ALL DATA SETS.
C      JMESS ARRAY IS USED TO THROW AWAY BAD RECORDS ON TAPE7.
      DIMENSION Y(2030),H(20),JMESS(20)
      DIMENSION CARD(20,100)
      CALL ERRSET (219,1,-1,1)
      JSET=0
      KSET=1
      NPSET=4
      READ (1,120) NDSET,NFSET,JMESS
120  FORMAT (I2,2X,I2,2X,20I2)
      IF (NFSET .GT. 0) GO TO 56
      NFSET=50
56  READ (1,122) RATE,HSCALE,HZERO,THETA,COILK
122  FORMAT (2F5.0,3F6.0)
      HSCALE=HSCALE/10000.
      THETA=THETA/100.
55  IF (NDSET .GE. NFSET) GO TO 50
      IEND=NDSET
      IF (IEND) 51,51,52
50  IEND=NFSET
52  DO 59 N=1,IEND
      JSET=JSET+1
      NPSET=NPSET+1
      IF (NPSET .EQ. 5) GO TO 45
      GO TO 40
45  WRITE (3,106)
106  FORMAT ('1')
      NPSET=1
40  READ (1,121,END=60) (CARD(K,JSET),K=1,20)
121  FORMAT (20A4)
5  IF(JSET.EQ.JMESS(KSET)) GO TO 53

```

```

1 READ (10,110,ERR=1000,EVD=57) (Y(I),I=1,1024)
110 FORMAT (32(32A4))
GO TO 3
1000 WRITE (3,111)
111 FORMAT ('      ERROR IN DATA TRANSFER FROM TAPE7')
CARD(10,JSET)=1111
READ (10,110) (Y(I),I=1,1024)
3 CALL CHGTC (Y)
CALL TRANS (Y,H)
WRITE (3,125) (CARD(K,JSET),K=1,20)
125 FORMAT ('0',20A4)
WRITE (3,123) RATE,HSCALE,HZERO,THETA,COILK
123 FCRMAT ('0 RATE=',F4.0,' KCPS. HSCALE=',F6.4,' HZERO=',F6.0
1,' GAUSS THETA=',F7.2,' DEG. COILK=',F6.0,' GAUSS/VOLT')
WRITE (9) (CARD(K,JSET),K=1,20)
WRITE (3,107)
107 FORMAT ('0 THE FIELD VALUES ARE: '/')
WRITE (3,103) (H(I),I=1,20)
103 FORMAT (10F12.2)
DO 6 J=1,20
6 H(J)=H(J)*(10./4096.)*COILK/HSCALE
WRITE (9) (Y(I),I=1,2008),(H(I),I=1,20),RATE,HSCALE,HZERO,
1THETA,CCILK
WRITE (3,108)
108 FCRMAT (' ')
WRITE (3,103) (H(I),I=1,20)
WRITE (3,104)
104 FORMAT ('0')
59 CONTINUE
END FILE 9
NDSET=NDSET-NFSET
GO TO 55
60 WRITE (3,130)
130 FORMAT ('0 YOU'RE OUT OF CARDS, BUDDY ')
GO TO 58
57 WRITE (3,101)
101 FORMAT ('0 CAN'T YOU EVEN COUNT--STUPID')

```

```

      JSET=JSET-1
      GO TO 58
53  KSET=KSET+1
      READ(10,110,ERR=1001,END=57)(Y(I),I=1,1024)
      GO TO 5
1001 WRITE(3,112)
      112 FORMAT('      ERROR IN READ AT NO. 53')
      READ(10,110)(Y(I),I=1,1024)
      GO TO 5
      51 WRITE (3,124)
124  FORMAT ('0      TRANSFER OF REQUESTED DATA SETS COMPLETED')
      58 REWIND 9
      REWIND 10
      WRITE (3,105)
105  FORMAT ('1      THE FOLLOWING ITEMS ARE CATALOGED: '/')
      IF (JSET .LE. 0) GO TO 19
      DO 54 J=1,JSET
      54 WRITE (3,102) (CARD(K,J),K=1,20)
102  FORMAT (20A4/)
      19 CONTINUE
      STOP
      END
//LKED.SYSIN DD *
      TRANS AND CHGTC OBJECT DECKS GO HERE....
//GO.FT10F001 DD UNIT=(TAPE7,,DEFER),VOLUME=SER=TP0132,LABEL=(1,NL), X      A
//      DISP=(OLD,KEEP), X      B
//      DCB=(DEN=1,TRTCH=ET,RECFM=F,BLKSIZE=4096,BUFNO=1) C
//GO.FT09F001 DD UNIT=(TAPE,,DEFER),VOLUME=SER=TP5028, X      01
//      DSN=SEQ.A0026WD1,LABEL=(1,SL), X
//      DISP=(NEW,KEEP) 03
//GO.FT03F001 DD SYSOUT=A,SPACE=(CYL,(2)), 04
//      DCB=(RECFM=FBA,LRECL=133,BLKSIZE=3325,BUFNO=1) 05
//GO.FT01F001 DD * 06
02  99
      250 6977      13565
      SEQ.A0026C50      2-050-001      THETA= -4.0 DEG.
      SEQ.A0026C50      2-050-002      THETA= -4.0 DEG.

```

```
//B334FFT JOB 'A0367,SI=224K,II=11,LI=6',JIM,MSGLEVEL=1
//STEP1 EXEC FORTH,PARM.FORT='MAP,DECK,OPT=2,LIST,XREF',
// REGION.FORT=224K,PARM.LKED='MAP,LIST,LET,HIAR',
// TIME.GO=8,REGION.GO=(160K,32K),LINES=6
//FORT.SYSUT2 DD UNIT=SPOOL,SPACE=(CYL,(1,1))
//FORT.SYSIN DD *
```

X
X

```
C
C DE HAAS-VAN ALPHEN FREQUENCY COMPUTATIONS
C *****THIS IS THE MAIN PROGRAM*****
C THIS PROGRAM CONTAINS THE FAST FOURIER FIT SUBROUTINE RHARM.
C
```

```
COMMON/AUTOF/FF(20),FL(20),NSTART(20),NSTOP(20)
COMMON/CALC/S(2048),INV(2048)
COMMON/DATA/H20S(20),Y1S(2008),HIS(2008)
COMMON/FIELD/DT,TZERO,THETA,HZERO,H20(20)
COMMON/FREQ/FFIRST(10),FLAST(10),TZERO1(10)
COMMON/GIANT/Y(16388)
COMMON/LABS/GLAB1(5),GLAB2(5),XLAB1(5),YLAB1(5),DDATE(7),
1RLAB(5),FLAB(5)
DIMENSION H(2008),HI(2008),Y1(2008)
EQUIVALENCE (H(1),HI(1),S(1)),(Y1(1),S(2009))
BPASS=0.0
OTMR=0.0
OTMC=0.0
MMM=13
MM=16384
424 CALL READIN (BPASS,OMEGA1,OMEGA2,STEP,STEPO,TMR,TMC,RES,MFSET,
1RATE,HSCALE,THETA1,COILK,NGRAPH,MSTART,MSTOP,ALIM,NTOUT,NSER)
KF=1
22 T000=TZERO1(KF)
TZERO=T000/(1.E06)
DO 30 I=1,20
30 H20(I)=H20S(I)
CALL BFIELD
CALL FITH (IFIELD,LRISE,LFALL)
DO 70 I=1,2008
HI(I)=1./H(I)
```



```

70 HIS(I)=H1(I)
   J=1
6  FF(J)=FFIRST(KF)
1  FL(J)=STEP*FF(J)
   IF (FL(J) .GE. FLAST(KF)) GO TO 2
   JJ=J+1
   FF(JJ)=STEPO*FF(J)
   J=JJ
   GO TO 1
2  FL(J)=FLAST(KF)
   IF (KF .GE. MFSET) GO TO 3
   KKF=KF+1
   IF (TZERO1(KKF) .NE. TGOO) GO TO 3
   J=J+1
   KF=KKF
   GO TO 6
3  IF (IFIELD .GT. 2) GO TO 4
   MFFCT=J
   NFTOT=J
   IF (IFIELD .LT. 2) GO TO 5
   DO 10 I=1,MFFCT
   JJ=I+MFFCT
   FF(JJ)=FF(I)
10  FL(JJ)=FL(I)
   NFICT=MFFCT+1
   NFFCT=2*MFFCT
   NFTOT=NFFCT
   GO TO 5
4  NFICT=1
   NFFCT=J
   NFTOT=J
5  IF (NFTOT .GT. 20) GO TO 26
   IF (IFIELD-2) 39,41,40
C  RISING FIELD ONLY.
39  LORISE=MSTART
   IF (MSTOP .GE. LRISE) GO TO 50
   LRISE=MSTOP

```

```

      GO TO 50
C     FALLING FIELD ONLY.
      40 LOFALL=MSTOP
        IF (MSTART .LE. LFALL) GO TO 50
        LFALL=MSTART
        GO TO 50
C     BOTH FIELDS.
      41 IF (MSTART .GE. LFALL) GO TO 43
        LORISE=MSTART
        IF (MSTOP .GE. LRISE) GO TO 42
        LRISE=MSTOP
        IFIELD=1
        GO TO 50
      43 LFALL=MSTART
        IFIELD=3
      42 LOFALL=MSTOP
        GO TO 50
      50 IF (IFIELD .GT. 2) GO TO 51
        IF ((LRISE-LORISE) .LE. 0) GO TO 28
        IF (IFIELD .LT. 2) GO TO 52
      51 IF ((LOFALL-LFALL) .LE. 0) GO TO 28
      52 IF (OMEGA1 .LE. 0.0) GO TO 53
        IF (OMEGA2 .LE. 0.0) GO TO 53
        IF (OMEGA2 .LE. OMEGA 1) GO TO 29
        CALL RANGE (IFIELD,1,MFFCT,NFICT,NFFCT,LORISE,LRISE,LFALL,LOFALL,
      1OMEGA1,OMEGA2,DT,RECYC)
        IF (RECYC .EQ. 1.0) GO TO 25
        GO TO 59
      53 IF (IFIELD .GT. 2) GO TO 55
        DO 54 I=1,MFFCT
          NSTART(I)=LORISE
      54 NSTOP(I)=LRISE
        IF (IFIELD .LT. 2) GO TO 59
      55 DO 56 I=NFICT,NFFCT
          NSTART(I)=LFALL
      56 NSTOP(I)=LOFALL
      59 ID=1

```

```

      GO TO 80
69  ID=IID
      DO 84 I=1,2008
        Y1(I)=Y1S(I)
84  HI(I)=HIS(I)
      GO TO 83
80  DO 81 I=1,2008
81  Y1(I)=Y1S(I)
83  CALL GENY (MM,ID,DRH,NPTS,RES,MF,ML,CZERO,DHI,ARES)
      IF (DRH .EQ. 0.) GO TO 27
      CALL RHARM (Y,MMM,INV,S,IFERR)
87  CALL DATE (DCATE)
      WRITE (3,102) GLAB1,GLAB2,DDATE
102  FCRMAT ('1  DE HAAS-VAN ALPHEN COMPUTATIONS--  DATA SET NUMBER
      1',10A4,3X,7A4)
      WRITE (3,103) RATE,HSCALE,HZERO,THETA1,COILK,TOOJ
103  FORMAT ('0  RATE=',F4.0,' KCPS.  HSCALE=',F6.4,'  HZERO=',F6.0
      1,' GAUSS  THETA=',F7.2,' DEG.  COILK=',F6.0,' GAUSS/VOLT  TZE
      1RO=',F4.0,' MSEC.')
      WRITE (3,109)
109  FORMAT ('0')
      OMEG1=OMEGA1/1000.
      OMEG2=OMEGA2/1000.
      WRITE (3,57) OMEG1,OMEG2
57  FORMAT ('  AMPLIFIER BANDWIDTH ASSUMED IS',F8.2,' KHZ  TO',
      $F8.2,' KHZ')
      WRITE (3,58) NSTART(ID),NSTOP(ID)
58  FORMAT ('  THE ANALYSIS STARTS AT COUNT NUMBER',I5,' AND EN
      IDS AT NUMBER',I5)
      IF (IFIELD .GT. 2) GO TO 90
      IF (ID .GT. MFFCT) GO TO 90
      WRITE (3,180) RLAB
180  FORMAT (80X,5A4)
      GO TO 95
90  WRITE (3,180) FLAB
95  CALL OUTPUT (ID,DRH,NPTS,NGRAPH,ARES,MF,ML,CZERO,DHI,ALIM,NTOUT,
      INSER)

```

```

      CALL STCPTM (TMR,TMC)
      TMRI=TMR-OTMR
      TMCI=TMC-OTMC
      OTMR=TMR
      OTMC=TMC
      CALL STARTM (TMR,TMC)
      WRITE (3,115) TMRI
115  FORMAT (70X,'REAL TIME USED IN EXECUTION OF DATASET (SEC)=',F7.2)
      WRITE (3,116) TMCI
116  FORMAT (70X,' CPU TIME USED IN EXECUTION OF DATASET (SEC)=',F7.2)
      96 IF (ID .GE. NFTOT) GO TO 24
         IID=ID+1
         IF (CZERO .NE. 1.0) GO TO 69
         IF (NSTART(IID) .NE. NSTART(ID)) GO TO 69
         IF (NSTCP(IID) .NE. NSTOP(ID)) GO TO 69
         ID=IID
         GO TO 87
      24 IF (KF .GE. MFSET) GO TO 424
         KF=KKF
         GO TO 22
      25 WRITE (3,100)
100  FORMAT ('0      ERROR ENCOUNTERED IN SUBROUTINE RANGE')
         GO TO 424
      27 WRITE (3,104)
104  FORMAT ('0      DRH WAS EQUAL TO ZERO')
         GO TO 424
      26 WRITE (3,101)
101  FORMAT ('0      PGM. TERMINATED---NFTOT EXCEEDED 20')
         GO TO 999
      28 WRITE (3,105) MSTART,MSTOP,LRISE,LFALL
105  FORMAT ('0      PGM. TERMINATED---MSTART=',I6,' ,MSTOP=',I6,
      1' ,LRISE=',I6,' ,LFALL=',I6)
         GO TO 999
      29 WRITE (3,106) CMEGA1,OMEGA2
106  FORMAT ('0      PGM. TERMINATED---OMEGA1=',E11.2,' ,OMEGA2=',E11.2)
999  STOP
      END

```

```

SUBROUTINE READIN (BPASS,OMEGA1,OMEGA2,STEP,STEPO,TMR,TMC,RES,
1MFSET,RATE,HSCALE,THETA1,COILK,NGRAPH,MSTART,MSTOP,ALIM,NTOUT,
1NSER)

```

C
C
C
C
C
C

```

THIS SUBROUTINE READS DATA FROM 9 TRACK TAPE.
THIS SUBROUTINE READS ALL CARD DATA FOR THIS PROGRAM.
A BLANK CARD INDICATES END OF FREQ. SET, OR SETS OF DATA.
MAX. OF NINE FREQ CARDS IN A SET.

```

```

COMMON/CALC/TCARD(10)
COMMON/DATA/H2OS(20),Y1S(2008)
COMMON/FIELD/DT,TZERO,THETA,HZERO,H2O(20)
COMMON/FREQ/FFIRST(10),FLAST(10),TZERO1(10)
COMMON/LABS/GLAB1(5),GLAB2(5),XLAB1(5),YLAB1(5),DDATE(7),
1RLAB(5),FLAB(5)
CALL ERRSET (219,1,-1,1)
REAL*4 BLANK/' '/
PIE=3.14159
IF (BPASS .GT. 0.0) GO TO 40
READ (1,101) XLAB1,YLAB1,RLAB,FLAB
101 FORMAT (20A4)
READ (1,110) EXTIME,OMEGA1,OMEGA2,STEP,STEPO,NGRAPH,NTOUT,NSER,
1ALIM
110 FORMAT (F5.1,2E10.2,2F5.2,3I5,F5.2)
IF (EXTIME .GT. 0.0) GO TO 80
EXTIME=10.0
80 EXTIME=EXTIME*60.0
TMR=0.0
TMC=0.0
CALL STARTM (TMR,TMC)
BFREQ=0.0
BPASS=1.0
71 K=1
85 READ (1,105,END=5) FFIRST(K),FLAST(K),TZERO1(K),RES,MSTART,MSTOP
C NOTE: RES, MSTART AND MSTOP MUST BE READ OFF THE 'BLANK' CARD
C SEPARATING FREQUENCY CARDS FROM DATA CARDS IF ANY VALUES ARE
C ARE TO BE RETAINED.

```

```

105 FORMAT (2E10.2,F5.0,F5.3,2I5)
    IF (FFIRST(K) .LE. 0.0) GO TO 11
    IF (FLAST(K) .LE. FFIRST(K)) GO TO 12
    K=K+1
    GO TO 85
11 MFSET=K-1
    IF (MSTART .NE. 0.0) GO TO 20
    MSTART=1
20 IF (MSTOP .NE. 0.0) GO TO 40
    MSTOP=2008
40 CALL STCPTM (TMR,TMC)
    IF (TMC .GT. EXTIME) GO TO 82
    CALL STARTM (TMR,TMC)
86 READ (1,106,END=5) TEND,CARDR,CARD
106 FORMAT (20X,A4,20X,2A4)
    IF (TEND .EQ. BLANK) GO TO 50
    REWIND 9
    GO TO 86
50 IF (CARDR .EQ. BLANK) GO TO 71
70 READ (9,ERR=1001,END=60) TCARD,GLAB1,GLAB2
    IF (GLAB1(2) .EQ. CARDR) GO TO 51
    READ (9,END=65)
    GO TO 70
51 IF (GLAB1(3) .EQ. CARD) GO TO 52
    READ (9,END=65)
    GO TO 70
52 READ (9,ERR=1002,END=61) Y1S,H2OS,RATE,HSCALE,HZERO,THETA1,COILK
    THETA=THETA1*PIE/180.
    IF (RATE .LE. 0.0) GO TO 7
    DT=1./(RATE*1000.)
    RETURN
7 WRITE (3,120)
120 FORMAT ('0 PROGRAM TERMINATED--RATE NOT .GT. ZERO')
    GO TO 84
12 WRITE (3,107)
107 FORMAT (' *****WATCH YOUR FREQUENCY CARDS--TOSS BAD CARD*****')
    GO TO 85

```

```

1001 WRITE (3,111)
111 FORMAT ('      ERROR IN TCARD TRANSFER AT STATEMENT 70')
    READ (9) TCARD,GLAB1,GLAB2
    WRITE (3,101) TCARD,GLAB1,GLAB2
    READ (9,END=63)
    GO TO 40
1002 WRITE (3,112)
112 FORMAT ('      ERROR IN DATA TRANSFER')
    WRITE (3,101) TCARD,GLAB1,GLAB2
    READ (9)
    GO TO 40
61 WRITE (3,113)
113 FORMAT ('      WHO GOOFED?  THE DATA SEEMS TO BE MISSING.')
    GO TO 40
63 WRITE (3,113)
    WRITE (3,101) TCARD,GLAB1,GLAB2
    GO TO 40
65 WRITE (3,113)
    WRITE (3,101) TCARD,GLAB1,GLAB2
60 GO TO 70
82 WRITE (3,451)
451 FORMAT ('1          END OF TIME    HOEKSTRA   A07 PHYS. TELE 4-6832')
    GO TO 84
5 WRITE (3,452)
452 FORMAT ('1          END OF DATA   HOEKSTRA   A07 PHYS. TELE 4-6832')
84 REWIND 9
    CALL STCPTM (TMR,TMC)
    CALL DATE (DDATE)
    WRITE (3,121) TMR
121 FORMAT (70X,'REAL TIME USED IN EXECUTION OF PROG (SEC)=',F7.2)
    WRITE (3,122) DDATE,TMC
122 FORMAT (42X,7A4,' CPU TIME USED IN EXECUTION OF PROG (SEC)=',F7.2)
    STOP
    END

```

```

      SUBROUTINE BFIELD
C
C      THIS SUBROUTINE CALCULATES THE 'B' FIELD FROM THE 20 'H' VALUES.
C      IT WILL CORRECT FOR BOTH TIME DELAY AND DEMAGNETIZING EFFECTS.
C
      COMMON/CALC/X1(20),W(20)
      COMMON/FIELD/DT,TZERO,THETA,HZERO,H20(20)
      DOUBLE PRECISION Q(6)
      X1(1)=42.
      W(1)=1.
      DO 1 I=2,20
      X1(I)=X1(I-1)+100.
1  W(I)=1.
      CALL CPLSPA (5,20,X1,H20,W,Q,0.0)
      DO 2 I=1,20
C      CALCULATE DERIVATIVE D=DH/DT
      X=X1(I)
      K=2
      D=Q(2)
9  D=D+K*Q(K+1)*X
      IF (K .GE. 5) GO TO 2
      K=K+1
      X=X*X1(I)
      GO TO 9
2  H20(I)=H20(I)-D*TZERO/DT
      IF (HZERO .LE. 0.0) GO TO 4
      P=(CCS(THETA))**2
      X=HZERO*(3.*P-1.)/2.
      Y=HZERO*P
      DO 3 I=1,20
3  H20(I)=HZERO+H20(I)*(H20(I)+X)/(H20(I)+Y)
4  RETURN
      END

```



```

SUBROUTINE FITH (IFIELD,LRISE,LFALL)
COMMON/CALC/H(2008),X1(20),Y1(3),W(20)
COMMON/FIELD/DUM(4),H20(20)
DCUBLE PRECISION Q(6)

C
C   IFIELD = 1  RISING FIELD
C             2  BOTH---ACROSS PEAK FIELD
C             3  FALLING FIELD
C   LRISE = STOPPING COUNT FOR RISING FIELD.
C   LFALL = STARTING COUNT FOR FALLING FIELD.
C
  IF (H20(1) .GE. H20(2)) GO TO 2
  I=2
1 IF (H20(I) .GE. H20(I+1)) GO TO 3
  I=I+1
  IF (I .LT. 20) GO TO 1
  IFIELD=1
  NUMB=2008
  I2=20
  GO TO 5
2 IFIELD=3
  NUMB=2008
  I2=20
  GO TO 5
3 DO 4 K=1,3
  W(K)=1.
  X1(K)=FLCAT(K-2)
  KK=I-2+K
4 Y1(K)=H20(KK)
  CALL OPLSPA (2,3,X1,Y1,W,Q,0.0)
  NN=-Q(2)*100./(2.*Q(3))
  NUMB=42.+100.*(I-1)+NN
  IFIELD=2
  I2=I
  IF (NN .GT. 0) GO TO 5
  I2=I-1
5 X1(1)=42.

```

```

      W(1)=1.
      DO 6 I=2,20
      X1(I)=X1(I-1)+100.
6     W(I)=1.
      IF (I2 .GT. 5) GO TO 7
      IFIELD=3
      DO 106 J=1,NUMB
106   H(J)=1.0
      GO TO 11
      7 CALL OPLSPA (5,I2,X1,H20,W,Q,0.0)
      N1=1
      N2=NUMB
      8 DO 10 I=N1,N2
      X=FLOAT(I)
      K=2
      H(I)=Q(1)
      9 H(I)=H(I)+Q(K)*X
      IF (K .GE. 6) GO TO 10
      K=K+1
      X=X*FLOAT(I)
      GO TO 9
10    CONTINUE
      IF (N1 .NE. 1) GO TO 14
      IF (IFIELD .NE. 2) GO TO 14
11    I3=20-I2
      IF (I3 .GE. 5) GO TO 12
      IFIELD=1
      N3=NUMB+1
      DO 111 J=N3,2008
111   H(J)=1.0
      GO TO 14
      12 DO 13 I=1,I3
      L=I+I2
      H20(I)=H20(L)
      13 X1(I)=X1(L)
      CALL OPLSPA (5,I3,X1,H20,W,Q,0.0)
      N1=NUMB+1

```

```
      N2=2008
      GO TO 8
14  IF (IFIELD .GT. 2) GO TO 17
      I=NUMB-100
15  II=I+1
      IF (H(I) .GE. H(II)) GO TO 16
      I=II
      IF (I-NUMB) 15,16,16
16  LRISE=I
      IF (IFIELD .LT. 2) GO TO 20
17  I=N1+100
18  II=I-1
      IF (H(II) .LE. H(I)) GO TO 19
      I=II
      IF (I-N1) 19,19,18
19  LFALL=I
20  RETURN
      END
```

```

SUBROUTINE RANGE (IFIELD,MFICT,MFFCT,NFICT,NFFCT,LORISE,LRISE,
1LFALL,LOFALL,OMEGA1,OMEGA2,DT,RECYC)
COMMON/AUTOF/FF(20),FL(20),NSTART(20),NSTOP(20)
CCOMMON/CALC/HI(2008)

```

```

C
C   FF(I) = INITIAL FREQ. ARRAY.   FL(I) = FINAL FREQ. ARRAY.
C   NSTART(I), NSTOP(I) ARE STARTING AND STOPPING COUNTS FOR I TH FREQ. RANGE.
C   RECYC IS SET = 0.0 IF NO ERROR, = 1.0 IF ERROR OCCURS.
C   IFIELD = 1 RISING FIELD,   = 2 BOTH,   = 3 FALLING FIELD.
C   MFICT,MFFCT ARE STARTING & STOPPING VALUES FOR SUBSCRIPT I FOR RISING FIELD.
C   NFICT,NFFCT ARE STARTING & STOPPING VALUES FOR SUBSCRIPT I FOR FALLING FIELD.
C   LORISE AND LRISE ARE COMPUTED START AND STOP COUNTS FOR RISING FIELD.
C   LFALL AND LOFALL ARE COMPUTED START AND STOP COUNTS FOR FALLING FIELD.
C   CMEGA1 AND OMEGA2 ARE LOW AND HIGH CUTOFF FREQUENCIES OF AMPLIFIER.
C   DT IS THE TIME INTERVAL.
C

```

```

      IF (IFIELD .GT. 2) GO TO 10
      I=MFICT
1   J=LORISE
2   OMEG=ABS(FF(I)*(HI(J+1)-HI(J))/DT)
      IF (OMEG .LE. OMEGA2) GO TO 3
      J=J+2
      IF (J-LRISE) 2,19,19
3   NSTART(I)=J
      I=I+1
      IF (I .GT. MFFCT) GO TO 4
      IF (FF(I)-FF(I-1)) 1,2,2
4   I=MFFCT
5   J=LRISE
6   OMEG=ABS(FL(I)*(HI(J)-HI(J-1))/DT)
      IF (OMEG .GE. CMEGA1) GO TO 7
      J=J-2
      IF (J-LORISE) 19,19,6
7   NSTOP(I)=J
      I=I-1
      IF (I .LT. MFICT) GO TO 10
      IF (FL(I)-FL(I+1)) 6,6,5

```

```

10 IF (IFIELD .LT.2) GO TO 18
    I=NFFCT
11 J=LFALL
12 CMEG=ABS(FL(I)*[HI(J+1)-HI(J)]/DT)
    IF (CMEG .GE. CMEGA1) GO TO 13
    J=J+2
    IF (J-LOFALL) 12,19,19
13 NSTART(I)=J
    I=I-1
    IF (I .LT. NFICT) GO TO 14
    IF (FL(I)-FL(I+1)) 12,12,11
14 I=NFICT
15 J=LOFALL
16 OMEG=ABS(FF(I)*[HI(J)-HI(J-1)]/DT)
    IF (OMEG .LE. CMEGA2) GO TO 17
    J=J-2
    IF (J-LFALL) 19,19,16
17 NSTOP(I)=J
    I=I+1
    IF (I .GT. NFFCT) GO TO 18
    IF (FF(I)-FF(I-1)) 15,16,16
18 RECYC=0.0
    GO TO 20
19 RECYC=1.0
20 RETURN
    END

```

```

SUBROUTINE GENY (MM, ID, DRH, NPTS, RES, MF, ML, CZERO, DHI, ARES)
DOUBLE PRECISION HH, DQH, A
COMMON/AUTOFF/FF(20), FL(20), NSTART(20), NSTOP(20)
CCMMCN/CALC/HI(2008), Y1(2008)
COMMCN/GIANT/Y(1)
LMIN=NSTART(ID)
LMAX=NSTOP(ID)
DHI=HI(LMIN)-HI(LMAX)
CZERC=1.0
IF (RES .EQ. 0.0) GO TO 1
ARES=RES
CFIRST=ABS(FF(ID)*DHI)
CRES=1./RES
IF (CFIRST .GE. CRES) GO TO 1
C   ADDED ZEROES NEEDED TO INCREASE RESOLUTION
MMM=MM*(CFIRST/CRES)
IF (MMM .GT. 6384) GO TO 61
MMM=6384
ARES=FLOAT(MMM)/(FLOAT(MM)*CFIRST)
61 DQH=DHI/(MMM-1)
   IF (DQH) 62, 681, 63
C   FALLING FIELD CASE
62 MF=1
   ML=MM-MMM
   LF=ML+1
   LL=MM
   GO TO 64
C   RISING FIELD CASE
63 LF=1
   LL=MMM
   MF=LL+1
   ML=MM
C   FILL PART OF Y ARRAY WITH ZEROES
64 DO 65 I=MF, ML
65 Y(I)=0.0
   CZERO=FLOAT(MMM)/FLOAT(MM)
   GO TO 2

```

```

1 DQH=DHI/(MM-1)
  LF=1
  LL=MM
2 IF (DQH) 682,681,683
681 WRITE (3,100)
100 FCRMAT ('          DQH=0")
  GO TO 1014
682 DQH=-DQH
  ASIN=-1.
  GO TO 1002
683 ASIN=+1.
1002 I=LMIN
1003 K=LF
1004 HH=HI(I)
1005 A=HI(I)
  34 G=HH-A
    Y(K)=Y1(I)+(Y1(I+1)-Y1(I))*G/(HI(I+1)-HI(I))
1006 K=K+1
1007 IF (K-LL) 35,35,50
  35 IF (ASIN) 39,39,36
  36 HH=HH-DQH
1008 G1=HH
1009 IF (G1-HI(I+1)) 40,34,34
  39 HH=HH+DQH
1010 G1=HH
1011 IF (G1-HI(I+1)) 34,34,40
  40 I=I+1
1012 A=HI(I)
1013 IF (I-LMAX) 41,50,50
  41 IF (HI(I+1)-HI(I)) 34,40,34
  50 NPTS=K-1
1014 DRH=DQH
69 RETURN
END

```

```

SUBROUTINE OUTPUT (ID,DRH,NPTS,NGRAPH,ARES,MF,ML,CZERO,DHI,ALIM,
INTOUT,NSER)
COMMON/AUTOF/FF(20),FL(20)
COMMON/CALC/Z(300),UU(300),Z1(150),UU1(150),Z1R(150)
COMMON/GIANT/Y(1)
COMMON/LABS/GLAB1(5),GLAB2(5),XLAB1(5),YLAB1(5),DDATE(7)
NZ1MAX=150
FACTOR=ABS(CZERO/DHI)
KFIRST=FF(ID)/FACTOR
KLAST=FL(ID)/FACTOR
WRITE (3,608) KFIRST,KLAST
608 FORMAT ('          KFIRST=',I6,'      KLAST=',I6)
IF (KFIRST .GE. KLAST) GO TO 461
WRITE (3,3333) DRH,NPTS,FACTOR
3333 FORMAT ('          DRH=',E16.7,'      NPTS=',I7,'      FACTOR=',E16.7)
IF (CZERO .EQ. 1.0) GO TO 10
RRES=ARES*100.
MFT=ML-MF+1
WRITE (3,110) MFT,MF,ML,RRES
110 FORMAT ('0',I13,' ZEROS WERE ADDED IN Y-ARRAY BETWEEN',I6,' AND',
1I6,' TO OBTAIN RESOLUTION OF',F6.3,' PER CENT.')
```

10 SCAN=299

```

510 IF (KLAST-KFIRST .GT. SCAN) GO TO 200
KSTART=KFIRST
KSTOP=KLAST
502 NFIRST=(KSTART*2)+1
NLAST=(KSTOP*2)+1
KFIRST=KFIRST+SCAN-10
J=0
DO 500 K=NFIRST,NLAST,2
J=J+1
Z(J)=(Y(K))**2 + (Y(K+1))**2
W=(K-1)/2
500 UU(J)=FACTOR*W
CUU1=UU(1)/1.E06
CUUJ=UU(J)/1.E06
WRITE (3,601) CUU1
```



```

601 FORMAT ('0          LOWEST FREQUENCY CONSIDERED WAS  =',F7.2,' MG')
   WRITE (3,602) CUUJ
602 FORMAT ('          HIGHEST FREQUENCY CONSIDERED WAS  =',F7.2,' MG')
   WRITE (3,5555) J
5555 FCRMAT ('          J IS =',I5)
      CALL PEAKS (J,Z,UU,LK,Z1,UU1,RATION,NZ1MAX)
      DO 501 I=1,J
501  Z(I)=Z(I)/RATION
      IF (NGRAPH .NE. 0) GO TO 499
426  CALL GRAPH (J,UU,Z,3,4,12.,10.,0,0,0.1,0.0,XLAB1,YLAB1,
1GLAB1,GLAB2)
499  IF (ALIM .GT. 0.0) GO TO 80
      JJ=0
      DO 802 I=1,LK
      Z1R(I)=Z1(I)/RATION
      IF (Z1R(I) .LT. 0.005) GO TO 802
      JJ=JJ+1
      Z1R(JJ)=Z1R(I)
      Z1(JJ)=Z1(I)
      UU1(JJ)=UU1(I)/1.E06
802  CONTINUE
      GO TO 428
80  JJ=0
      DO 902 I=1,LK
      IF (Z1(I) .LT. ALIM) GO TO 902
      JJ=JJ+1
      Z1R(JJ)=Z1(I)/RATION
      Z1(JJ)=Z1(I)
      UU1(JJ)=UU1(I)/1.E06
902  CONTINUE
428  WRITE (3,415) RATION
415  FORMAT ('          NORMALIZING FACTOR (FAST FOURIER FIT)  =',F8.2)
      WRITE (3,1001) NSER
1001 FCRMAT ('0          FAST FOURIER FIT ANALYSIS',30X,'SERIAL NO. =',I5)
      WRITE (3,417)
417  FCRMAT ('0          INTENSITY          FREQUENCY (MG)          NORM. INTENS
      IITY')

```

```

      WRITE (3,419)
419  FORMAT (' ')
      IF (JJ .EQ. 0) GO TO 1070
820  DO 420 I=1,JJ
420  WRITE (3,1006) Z1(I),UU1(I),Z1R(I)
1006 FORMAT (3F18.3)
      IF (NTOUT .EQ. 0) GO TO 1070
      WRITE (8) NSER,GLAB1,GLAB2,DDATE,CUU1,CUUJ,JJ
      WRITE (8) ((Z1(I),I=1,JJ),(UU1(K),K=1,JJ))
      NSER=NSER+1
1070 WRITE (3,416)
416  FORMAT ('0')
      IF (KFIRST - KLAST) 510,461,461
200  KSTART=KFIRST
      KSTOP=KSTART+SCAN
      GO TO 502
461  RETURN
      END

```

```

      SUBROUTINE PEAKS (NUM,Z,UU,NUM1,Z1,UU1,RATION,NZ1MAX)
C
C   THIS SUBROUTINE FINDS THE PEAKS IN THE SPECTRUM
C   THE Z,UU,Z1,UU1 ARRAYS MUST BE DIMENSIONED IN THE CALLING PGM.
C   NUM1=# OF POINTS RETURNED BY Z1 OR UU1 ARRAYS.
C   NUM=# OF POINTS SUPPLIED IN Z OR UU ARRAYS.
C   RATION = LARGEST VALUE OF Z OR Z1 ARRAYS.
C   NZ1MAX=SIZE OF Z1 ARRAY AS DIMENSIONED BY CALLING PGM.
C       Z = AMPLITUDE OF DATA      UU = FREQ OF CORRESPONDING DATA.
C       Z1 = HEIGHT OF PEAKS      UU1 = FREQ OF CORRESPONDING PEAKS.
C
      DIMENSION Z(1),UU(1),Z1(1),UU1(1)
      DIMENSION X1(3),X2(3),W(3)
      DOUBLE PRECISION A(4)
      Z1(1)=0.0
      K=0
      I=1
      IF (Z(2)-Z(1)) 405,407,401
407 I=2
401 IF (Z(I+1)-Z(I)) 403,403,402
402 I=I+1
      IF (I-NUM) 401,410,410
405 IF (Z(I+1)-Z(I)) 406,401,401
406 I=I+1
      IF (I-NUM) 405,410,410
403 DO 404 J=1,3
      NM=I+J-2
      W(J)=1.0
      X1(J)=UU(NM)
404 X2(J)=Z(NM)
      CALL OPLSPA (2,3,X1,X2,W,A,0.0)
      K=K+1
      UU1(K)=-A(2)/(2.*A(3))
      Z1(K)=A(1)-(A(2)*A(2))/(4.*A(3))
      IF (K-NZ1MAX) 405,844,844
844 WRITE (3,846) NZ1MAX
846 FORMAT ('0 ***** THERE ARE MORE THAN ',I3,' PEAKS *****')

```

```

410 NUM1 = K
    IF (NUM1) 430,430,432
430 WRITE (3,431)
431 FORMAT ('0          THERE ARE NO PEAKS IN THIS DATA SET')
432 RATION=Z(1)
    IF (Z(NUM) .LE. RATION) GO TO 412
    RATION=Z(NUM)
412 DO 413 K=1,NUM1
    IF (Z1(K) .LE. RATION) GO TO 413
    RATION=Z1(K)
413 CONTINUE
414 RETURN
END
//LKED.SYSIN DD *
H1ARCHY 1,AUTOF,DATA,LABS
//GO.FT14F001 DD DSNAME=&SM,UNIT=SPOOL,DISP=(NEW,PASS),
// SPACE=(800,(120,15)),DCB=(RECFM=VS,LRECL=796,BLKSIZE=800)
//GO.FT09F001 DD UNIT=(TAPE,,DEFER),
// DSNAME=SEQ.A0026WD1,LABEL=(1,SL),
// DISP=(OLD,KEEP)
//GO.FT08F001 DD UNIT=(TAPE,,DEFER),
// DSNAME=SEQ.A0026TDA,LABEL=(1,SL),
// DISP=(NEW,KEEP)
//GO.FT03F001 DD SYSOUT=A,SPACE=(CYL,(5)),
// DCB=(RECFM=FBA,LRECL=133,BLKSIZE=3325,BUFNO=1)
//GO.FT01F001 DD *
DHVA FREQUENCY      INTENSITY      RISING FIELD      FALLING FIELD
10.0 2.50E C3 1.00E 05 2.40 2.30 1 1 1 0.80
4.10E C6 9.00E 06 40
2.00E 07 3.50E 07 40
*****THIS CARD IS BLANK*****
SEQ.A0026C50 2-050-001 THETA= -4.0 DEG.
SEQ.A0026C50 2-050-002 THETA= -4.0 DEG.
//SMPLTTR EXEC PLOT,PLOTTER=INCRMNTL

```

SMPLT1/3

SMPLT2/3

X 01

X

03

X 01

X

X

COMMENTS ON LIBRARY PROGRAMS CALLED:

DATE

DIMENSION X(7)

CALL DATE (X)

RESULTS IN CURRENT DATE AND TIME IN 7A4 FORMAT.

STARTM & STOPTH

A = 0.0

B = 0.0

CALL STARTM (A,B)

** OTHER PROGRAM STATEMENTS

CALL STOPTH (A,B)

RESULTS IN A = REAL TIME ELAPSED BETWEEN CALLS TO STARTM & STOPTH.

RESULTS IN B = CPU TIME ELAPSED BETWEEN CALLS TO STARTM & STOPTH.

OPLSPA

OPLSPA IS A ROUTINE TO LEAST SQUARES FIT A POLYNOMIAL
OF DESIRED DEGREE TO GIVEN DATA.

DOUBLE PRECISION Q (NDEG+1)

CALL OPLSPA (NDEG,NPTS,X,Y,W,Q,0.0)

NDEG = DEGREE OF DESIRED POLYNOMIAL = MAX. VALUE OF 10.

NPTS = NUMBER OF POINTS.

X = FIRST ELEMENT OF TABLE OF INDEPENDENT VARIABLES.

Y = FIRST ELEMENT OF TABLE OF DEPENDENT VARIABLES.

W = FIRST ELEMENT OF TABLE OF WEIGHTS ASSOCIATED WITH THE
CORRESPONDING DATA POINT X(I),Y(I). IF DATA IS UNWEIGHTED,
SET W(I) = 1.0 FOR I = 1 TO NPTS.

Q = FIRST ELEMENT OF TABLE OF COEFFICIENTS RETURNED BY OPLSPA.
TABLE ARRANGED TO CORRESPOND TO ASCENDING POWERS OF
THE INDEPENDENT VARIABLE.